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Faint Standards for *ZYJHK* from the UKIDSS and VISTA Surveys

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ABSTRACT

The currently defined “UKIRT Faint Standards” have *JHK* magnitudes between 10 and 15, with $K_{\text{median}} = 11.2$. These stars will be too bright for the next generation of large telescopes. We have used multi-epoch observations taken as part of the UKIRT Infrared Deep Sky Survey (UKIDSS) and the Visible and Infrared Survey Telescope for Astronomy (VISTA) surveys to identify non-variable stars with *JHK* magnitudes in the range 16 – 19. The stars were selected from the UKIDSS Deep Extragalactic Survey (DXS) and Ultra Deep Survey (UDS), the WFCAM calibration data (WFCAM-CAL08B), the VISTA Deep Extragalactic Observations (VIDEO) and UltraVISTA. Sources selected from the near-infrared databases were paired with the Pan-STARRS Data Release 2 of optical to near-infrared photometry and the Gaia astrometric Data Release 2. Colour indices and other measurements were used to exclude sources that did not appear to be simple single stars. From an initial selection of 169 sources, we present a final sample of 81 standard stars with *ZYJHK* magnitudes, or a subset, each with 20 to 600 observations in each filter. The new standards have $Ks_{\text{median}} = 17.5$. The relative photometric uncertainty for the sample is < 0.006 mag and the absolute uncertainty is estimated to be $\lesssim 0.02$ mag. The sources are distributed equatorially and are accessible from both hemispheres.

Key words: standards – methods observational – techniques photometric

1 INTRODUCTION

Optical and infrared sky surveys have produced data with excellent astrometric and photometric precision. In 2019, we are benefiting from data releases by ground- and space-based surveys that were many years in the planning.

The Gaia mission (Gaia Collaboration et al. 2016) issued Data Release 2 on 2018 April 25. This release contains astrometric results for about 1.7 billion stars brighter than optical magnitude 21; parallaxes and proper motions are given for about 1.3 billion of these (Gaia Collaboration et al. 2018). The Sloan Digital Sky Survey (SDSS, York et al. 2000) and the Panoramic Survey Telescope & Rapid Response System (Pan-STARRS, Chambers et al. 2016) have imaged large areas of the Northern sky in blue/green to far-red/near-infrared filters. Tonry et al. (2012) give transformations between the SDSS and Pan-STARRS photometric systems. After transformation, the rms difference between the SDSS and Pan-STARRS *griz* photometry is 8 mmag, following a recalibration of the SDSS photome-

try with new flat fields and zero points derived from Pan-STARRS (Finkbeiner et al. 2016). This is consistent with Padmanabhan et al. (2008) who report a relative calibration accuracy of 0.7 – 1.3% for each filter in the SDSS.

Several near-infrared surveys of the sky have been undertaken in the last two decades. The 2-micron All Sky Survey (2MASS, Skrutskie et al. 2006) was executed between 1997 and 2001, with 1.3 m telescopes in both hemispheres providing complete sky coverage. The UKIRT Infrared Deep Sky Survey (UKIDSS, Lawrence et al. 2007) consisted of several Northern-hemisphere sub-surveys using the WFCAM camera and *ZYJHK* filters, and was executed between 2007 and 2011 on the 3.8 m UKIRT on Mauna Kea. Following UKIDSS, in 2012, the UKIRT Hemisphere Survey (UHS, Dye et al. 2018) began, and aims to provide continuous coverage in the *J* and *K* bands over the Declination range of zero to +60 degrees. A collaboration between the University of Hawaii and the United States Naval Observatory is continuing the UKIRT survey and adding the *H* band (Hodapp et al. 2018). In the Southern hemisphere, the 4.1 m Visible and Infrared Survey Telescope for Astronomy (VISTA Sutherland et al. 2015) started surveying the sky

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in 2009; the first surveys are complete or nearing completion as of 2019¹; VISTA is producing several sub-surveys using the *ZYJHKs* filters, or a subset. [González-Fernández et al. \(2018\)](#) compare the VISTA and UKIDSS photometric systems using equatorial stars observed with both cameras. The differences are small, and after transformation the rms difference between the measurement sets is 1 – 3 mmag at *JHK*.

The near-infrared surveys cover the entire sky and provide 1 – 10 sources per square arcminute on average. Such stars could be used to calibrate science data to ~10% ([González-Fernández et al. 2018](#); [Hodgkin et al. 2009](#)). Photometric standard stars are needed to calibrate data more accurately. In the near-infrared these have frequently been provided by UKIRT measurements published by [Hawarden et al. \(2001\)](#) and [Leggett et al. \(2006\)](#). The UKIRT Faint Standards (FS) cover a range in Declination of –25 to +55 degrees. They have *JHK* Vega magnitudes of 10 – 15 with $K_{\text{median}} = 11.2$ and $\sigma_{\text{median}} = 0.01$. The UKIRT FS require integration times of 1 – 10 seconds when observed on 8 m telescopes such as at the Gemini Observatory. On future telescopes with diameters three or more times larger, the integration times will be < 1 s; such short integrations are not only inefficient, they may give rise to calibration problems such as poor linearity corrections (e.g. [Leggett et al. 2006](#)).

In this paper we identify stars in the UKIDSS and VISTA surveys that are fainter than the UKIRT FS by a factor of ~100. We select sources that have a large number of repeat measurements which show them to be non-variable, and which have high precision photometric measurements. We pair the list of candidates with the Pan-STARRS and Gaia databases and use the optical and near-infrared colours to refine the sample and produce a set of well-measured and well-behaved single stars.

2 THE SURVEYS

The Northern hemisphere UKIDSS (with the WFCAM camera) and the Southern hemisphere VISTA use similar-size telescopes, and similar filters defined according to the Mauna Kea Observatories filter specifications ([Tokunaga et al. 2002](#)). Survey data generated by both telescopes are initially processed by the Cambridge Astronomy Survey Unit (CASU) and then transferred to the Wide-Field Astronomy Unit (WFAU) in Edinburgh for further processing and archiving ([Cross et al. 2012](#); [Hambly et al. 2008](#)). Table 1 gives a summary of the properties of the surveys used in this work.

Each of the surveys has a source table (e.g. `dxsSource`) which contains band-merged detections from deep images. In addition, for the surveys used here, there is a variability table (e.g. `dxsVariability`) which contains the light-curve statistics for each primary source that is detected in at least one epoch. Details of the multi-epoch table structure and processing can be found in [Cross et al. \(2009\)](#). The number of good observations for each filter and each source is given as `nGoodObs` in the variability table; these are measurements

Table 1. Overview of the Sky Survey Data

Survey Name	Area deg ²	Filters	Vega mag 5 σ Limits	Ref.
UKIDSS Deep Extragalactic Survey (DXS)	35.0	<i>JHK</i>	21 – 22	1
UKIDSS Ultra Deep Survey (UDS) ^a	0.8	<i>JHK</i>	23 – 25	1
VISTA Deep Extragalactic Observations (VIDEO)	12.0	<i>ZYJHKs</i>	24 – 25	2
VISTA UltraVISTA ^b	1.5	<i>YJHKs</i>	22 – 24	3
WFCAM Calibration WFCAMCAL08	10.4	<i>ZYJHK</i>	18 – 19	4

References: (1) [Lawrence et al. \(2007\)](#), (2) [Jarvis et al. \(2013\)](#), (3) [McCracken et al. \(2012\)](#), (4) [Ferreira Lopes et al. \(2015\)](#).

^a For the UDS, the source table aperture magnitudes are not aperture corrected and the variability table mean magnitudes are. For a sample of 30 objects, we find an average offset in photometry of 0.190 ± 0.005 , 0.202 ± 0.006 , and 0.179 ± 0.006 , at *J*, *H* and *K* respectively, such that the mean magnitudes are brighter.

^b For UltraVISTA there are differences in the aperture corrections and zeropoints used for the source table aperture magnitudes and the variability table mean magnitudes. For a sample of 19 objects we find $\text{meanMag} - \text{aperMag} = 0.020 \pm 0.005$, 0.052 ± 0.006 , -0.025 ± 0.010 , and -0.004 ± 0.012 at *Y*, *J*, *H* and *Ks* respectively.

which are not flagged as blended, saturated, etc. using the post-processing error bit flag (`ppErrBits`), i.e. a good observation has `ppErrBits=0`². The source and variability tables share the same primary key `sourceID`.

Our goal is to identify stars that are not variable, that have well-measured *JHK* magnitudes, and are fainter than $K = 16$. Therefore we only used surveys with variability tables for all of the *J*, *H*, and *K* filters. This means we include the UKIDSS Deep Extragalactic Survey (DXS), Ultra Deep Survey (UDS) and WFCAMCAL calibration data, but do not include the UKIDSS Galactic Clusters, Galactic Plane or Large Area Surveys (GCS, GPS, LAS). Similarly we use the VISTA Deep Extragalactic Observations Survey (VIDEO) and UltraVISTA surveys, but do not include the VISTA Kilo-Degree Infrared Galaxy, Magellanic Clouds, Variables in the Via Lactea, or Hemisphere Surveys (VIKING, VMC, VVV, VHS). The DXS, UDS, WFCAMCAL and VIDEO surveys are complete. The UltraVISTA survey is being continued³ with a final data release planned for 2021⁴.

The variability flag for each source (`variableClass`) is determined from repeat measurements and is stored in the variability table ([Cross et al. 2009](#)). The value is determined by the significance of the weighted average of the intrinsic noise over the expected noise, as given by a noise model for each pointing and each band. The weighting of the intrinsic noise is based on the number of good observations: $w_f = \frac{N_{\text{obs},f} - N_{\text{min}}}{N_{\text{obs},\text{max}} - N_{\text{min}}}$, where $N_{\text{min}} = 5$ is the minimum number of observations necessary to measure variability, $N_{\text{obs},f}$ is the number of good observations in that

¹ <https://www.eso.org/sci/observing/PublicSurveys/sciencePublicSurveys.html>

² <http://wsa.roe.ac.uk/ppErrBits.html>

³ <https://www.eso.org/sci/observing/PublicSurveys/sciencePublicSurveys.html#vistacycle2>

⁴ https://www.eso.org/sci/observing/PublicSurveys/docs/UltraVISTA2_SMP_03022017.pdf

band, and $N_{obs,max}$ is the maximum number of good observations for the star in any band. The default classification is non-variable (`variableClass` = 0), and if a star has fewer than 5 good observations in all bands it will be classified as non-variable; none of the objects selected here fall into this category. If a star has fewer than 5 observations in all bands but one, that band will determine the classification. If there are hundreds of observations in one band and tens in the others, the band with hundreds will have a strong weighting compared to the others. The star is classified as variable (`variableClass` = 1) if the ratio of the weighted intrinsic noise to the expected noise is > 3 .

We adopt the mean photometric magnitudes and uncertainties given in the variability tables as the reference calibration data in this work, and not the source table aperture magnitudes, for the following reasons. In some cases, the source tables do not go as deep as the stacked variability images; this is the case for the WFCAMCAL08 database where it was important to avoid source blending for calibration purposes. For the Ultra Deep Survey the source table photometry `AperMag` is not aperture-corrected, while the variability photometry `MeanMag` is (Table 1, Section 3). For the most recently processed survey, UltraVISTA, there are small differences between the variability and source table photometry due to different aperture corrections and Vega-to-AB zeropoint corrections (Table 1, Section 3). These issues will be corrected in a forthcoming UltraVISTA release. The variability photometry is processed in the same way for all the surveys used here, and so provides a self-consistent sample.

3 THE SQL SELECTIONS

We used the WFCAM Science Archive⁵ and the VISTA Science Archive⁶ to query the DXS, UDS, UltraVISTA, VIDEO and WFCAMCAL databases. The most recent data releases available at the time of writing were used: Data Release 11 of the DXS and UDS was used, Data Release 4 for UltraVISTA, Data Release 5 for VIDEO, and WFCAMCAL08B. For all surveys we selected for the photometric source to be the same as the source in the variability catalogue and for non-variable sources, for example `SELECT FROM dxsSource AS s, dxsVariability AS v WHERE s.sourceID=v.sourceID AND v.variableClass=0`.

We selected for point sources using the surveys' morphological classifications. For the shallower surveys – DXS and WFCAMCAL – the classification scheme uses a statistic which describes how point-like each object is with respect to an empirically derived, idealized radial profile representing the point source function (PSF) for the frame (Irwin et al. 2004; Hambly et al. 2008). The deeper surveys – UDS, UltraVISTA and VIDEO – use the TERAPIX SWARP image resampling tool (Bertin et al. 2002) and the `CLASS_STAR` statistic⁷ generated by the SEXTRACTOR software (Bertin & Arnouts 1996), together with magnitude cuts in each band (Liske et al. 2003; Warren et al.

2007). We found that we could select a good-sized sample of objects (more than ten) from each of the DXS, UDS, and WFCAMCAL surveys by selecting for stars only, using `s.mergedClass=-1`. For VIDEO we relaxed the selection to stars and probable stars, with `s.mergedClass` in $(-1, -2)$, in order to get a useful sample. For UltraVISTA we found that the `mergedClass` statistic in the source table classifies all sources with multi-filter photometry as galaxies. A visual inspection of the sources in the stacked images showed that the PSFs in the *Y* and *J* images are extended. We expect that future releases of the ongoing UltraVISTA survey will include improved morphological classifications. For this work, for UltraVISTA, we selected for possible stars by using the class statistic determined from the *H* and *K* images, `s.hclassStat>0.7 AND s.kscsclassStat>0.7`, and we use additional profile and colour selections to separate galaxies from stars (Section 6).

We selected for precise photometry by restricting the uncertainty in the mean magnitude to ≤ 0.006 mag. The photometric uncertainties are given in the variability catalogue by the rms value of the multiple measurements (`MagRms`) and the median absolute deviation of the magnitude (`MagMAD`): we selected for sources where these values scaled as expected with `nGoodObs`, for example `(v.jMagMAD/SQRT(v.jnGoodObs - 1))<=0.004 AND (v.jMagRms/SQRT(v.jnGoodObs - 1))<=0.006`.

For the DXS, VIDEO and WFCAMCAL surveys we selected for consistency between the source and variability table magnitudes by limiting the difference to 2.5σ where σ is determined from the variability rms and the source aperture magnitude error, for example, `(s.jAperMag3 - v.jMeanMag) < 2.5 * (SQRT(s.jAperMag3err * s.jAperMag3err + (v.jMagRms/SQRT(v.jnGoodObs - 1)) * (v.jMagRms/SQRT(v.jnGoodObs - 1))))`. For the UDS and UltraVISTA surveys, where there are systematic differences between the source and variability table magnitudes (Section 2 and Table 1), we selected for objects with a small range around the average offset; for example for the UDS we used `(s.jAperMag3 - v.jMeanMag < 0.20) AND (s.jAperMag3 - v.jMeanMag > 0.18)`, and for UltraVISTA we used `((s.jAperMag3 - v.jMeanMag) <= -0.04) AND ((s.jAperMag3 - v.jMeanMag) >= -0.08)`.

A limit on target declination was also implemented. In Sections 6 and 7 we use Pan-STARRS optical data, together with the near-infrared data, to further refine the star/galaxy separation and to remove sources which may be multiple, as evidenced by unusual colours. For this reason we restricted our searches to declination $> -30^\circ$.

The constraints on brightness and number of observations in each filter varied with each survey. Sources with $K > 16$ or $K > 16.5$ were selected from the DXS, UltraVISTA and WFCAMCAL, while sources with $K > 17.5$ were selected from the deeper UDS and VIDEO surveys (Table 1). For the data to be of calibration quality, we adopt a minimum number of measurements of 20 for each object in each filter. Less than twenty *H*-band observations were obtained for some objects in the DXS, and for that survey we restricted the search to a minimum of 20 observations in *J* and *K* and 5 in *H*, in order to use the colour information. The minimum number of observations for the other surveys ranged from 20 to 100 for each filter. While the *JHK* filters

⁵ <http://wsa.roe.ac.uk>

⁶ <http://surveys.roe.ac.uk/vsa>

⁷ <https://sextractor.readthedocs.io/en/latest/ClassStar.html>

are our priority in this work, we also record the Z and Y magnitudes where available.

The SQL queries used for each survey in their complete form are given in the Appendix. The searches produced 169 sources: 34 from the DXS, 51 from the UDS, 25 from UltraVISTA, 47 from the VIDEO survey, and 12 from WFCAM-CAL.

4 THE WFCAM-UKIDSS AND VISTA PHOTOMETRIC SYSTEMS

Both WFCAM and VISTA use filters specified by the Mauna Kea Observatories system (Tokunaga et al. 2002). There are small differences in the filters as delivered, and differences in the site and telescope optics, which lead to small differences between the native UKIDSS and VISTA photometric systems. González-Fernández et al. (2018) use a large sample of reddening-free stars with photometric errors < 0.1 mag, observed with both cameras, to derive the following color transformations between VISTA and WFCAM:

$$Z_V - Z_W = -(0.037 \pm 0.008) \times (J - K)_W + (0.040 \pm 0.005)$$

$$Y_V - Y_W = -(0.010 \pm 0.003) \times (J - K)_W - (0.048 \pm 0.002)$$

$$J_V - J_W = -(0.028 \pm 0.002) \times (J - K)_W - (0.004 \pm 0.001)$$

$$H_V - H_W = -(0.037 \pm 0.001) \times (J - K)_W + (0.025 \pm 0.001)$$

$$K_{SV} - K_W = (0.017 \pm 0.003) \times (J - K)_W - (0.022 \pm 0.002)$$

VISTA survey observations are continuing, and the photometry presented here can be enhanced by future UltraVISTA data releases, hence we convert the UKIDSS photometry to the VISTA system and use that as our reference system in this work.

The UKIDSS UDS field (Data Release 11) and the VISTA VIDEO XMM-Newton field (Data Release 5) overlap by 0.1 deg in Right Ascension and 0.9 deg in Declination. A search for non-variable point sources in the common region with $\text{AperMag3err} \leq 0.006$ produced a sample of 43 stars with H and K magnitudes between 13 and 16 (no J measurements were available). Allowing for small color transformations (see above), we find a mean difference between the UDS and VIDEO survey magnitudes of 0.010 mag at H and K . These results suggest that the color-transformed standard star photometry presented in this paper is robust at the 1% level.

5 MATCHING TO PAN-STARRS AND GAIA

The coordinates of the sources produced by our SQL searches of the WFCAM and VISTA science archives, described above, were matched to Data Release 2 of the Pan-STARRS data archive⁸ and Data Release 2 of the Gaia data archive⁹, which were the most recent releases at the time of

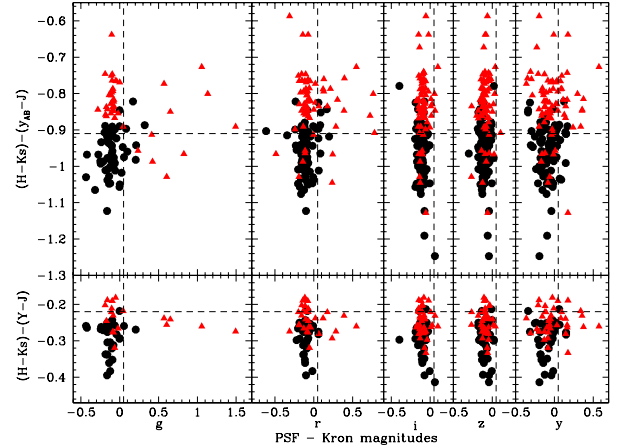


Figure 1. Red triangles are identified as possible galaxies either by the PSF – Kron magnitude or the colour index. The limits on these values are indicated by the dashed lines, see Section 6.

writing. The Pan-STARRS *grizy* magnitudes ranged from 18 to 23.

For about one-third of the sample, one or two of the Pan-STARRS filters have an aperture and PSF magnitude that differ by more than 2.5σ . About 60% of these we identify as possible galaxies below. The majority of the remainder have PSF and aperture magnitude differences that are not large in absolute terms, and it is possible that the uncertainties are slightly underestimated in this Pan-STARRS release.

Sources with Pan-STARRS AB magnitudes $r \lesssim 21$ and $i \lesssim 20.5$ were detected by Gaia, consistent with the $G < 21$ limit of Gaia Data Release 2 (Gaia Collaboration et al. 2018). Of those, about 85% have proper motion measurements (all $\lesssim 30$ mas yr⁻¹). About one-third of the sources with proper motion measurements have a trigonometric parallax measurement that is positive and has an uncertainty smaller than the parallax measurement. Fifteen of the 169 objects found in our searches have a Gaia parallax measurement that is significant. The next release of Gaia data, expected in the third quarter of 2020, will include additional and improved astrometry¹⁰.

6 STAR/GALAXY SEPARATION

Our goal here is to provide as clean a sample as possible of truly point source calibrators, so that, even if imaged at very high resolution, the aperture corresponding to the photometry is unambiguous. We therefore further prune the sample by excluding possible galaxies.

Davies et al. (2018) determined that stars and galaxies can be separated in VISTA data by using a near-infrared colour index. Davies et al. (2018) select for galaxies by applying the colour cut $(H - Ks) - (Y - J) > -0.26$, and verified the selection by visual inspection of objects brighter than $Y = 21.2$. Not all of our sample has a Y -band measurement and we determined the $(H - Ks) - (y_{AB} - J)$ color, replacing the VISTA Y_{Vega} with the Pan-STARRS y_{AB} magni-

⁸ <https://catalogs.mast.stsci.edu/>

⁹ <https://gea.esac.esa.int/archive/>

¹⁰ <https://www.cosmos.esa.int/web/gaia/release>

tude. Calibrating that index against the VISTA index, we adopt as a stellar indicator $(H - Ks) - (Y - J) < -0.22$ or $(H - Ks) - (y_{AB} - J) < -0.91$.

We also explored the use of the Pan-STARRS PSF and Kron magnitudes, searching for the brighter Kron magnitudes that would be expected if the source was extended (e.g. Chambers et al. 2016, their Figure 19). Chambers et al. (2016) find that stars can be selected by $\text{PSF} - \text{Kron} < 0.05$ magnitude, for magnitudes as faint as ~ 21 .

Figure 1 combines these two indicators, plotting $(H - Ks) - (Y - J)$ and $(H - Ks) - (y_{AB} - J)$ against $\text{PSF} - \text{Kron}$ magnitudes for each of the Pan-STARRS filters. The star/galaxy cuts are shown in the Figure; we excluded as possible galaxies sources which are either too red in one or both of the color indices, or too bright in any of the $\text{PSF} - \text{Kron}$ colours. The sources that remain are shown in black in Figure 1, and lie to the lower left, or have error bars that would place them in the lower left, of each panel (error bars are omitted from the Figure for clarity). These cuts identified 77 of the 169 sources as possible galaxies, and we omit them from the standard star sample. The Appendix Table B1 lists the possible galaxies together with Pan-STARRS and VISTA-system photometry, and Gaia data where available.

7 EXCLUSION OF MULTIPLE SYSTEMS AND OTHER CONTAMINANTS

We further refine the likely-star sample of 92 objects by excluding sources with atypical colours. These may be multiple systems, or the colours may be compromised in some way. To produce a conservative sample, we reject outliers from colour sequences prescribed by the majority of the sample. Figure 2 shows bluer colors — $g_{AB} - r_{AB}$, $r_{AB} - i_{AB}$, $i_{AB} - z_{AB}$ — and Figure 3 redder colours — $z_{AB} - y_{AB}$, $y_{AB} - J_{Vega}$, $J_{Vega} - H_{Vega}$. We identify 7 sources as outliers; these are shown as open circles in Figures 2 and 3 and listed in Appendix Table C1.

Figures 2 and 3 include color sequences produced by stellar model atmospheres. We used the BT-SETTL models (Allard et al. 2012; Baraffe et al. 2015) and the PARSEC models (Bressan et al. 2012; Girardi et al. 2000), for a range of metallicity as described in the Figure caption. For both model sets the $g - r$ colour deviates significantly from observation for M-type stars; this may be due to some issue common to the model atmospheres or it may be due to errors in the g bandpass adopted by both modelling groups, which could introduce a colour term. The colors and absolute magnitudes (where available), combined with the models, imply that the sample consists of dwarf stars with masses between 0.1 and 1.0 M_{\odot} .

Lastly, for the DXS, VIDEO and WFCAMCAL surveys, where there is no offset between the source table aperture magnitudes and the variability table mean magnitudes (Sections 2, 3), we excluded sources where these values deviated significantly. The earlier selections excluded discrepant sources based on the variability rms, now we exclude sources based on the variability MAD values. We excluded sources where the aperture and mean magnitudes differed by > 0.03 mag and the difference was significant by $> 2.5 \sigma$. Four additional objects were rejected by this criterion and

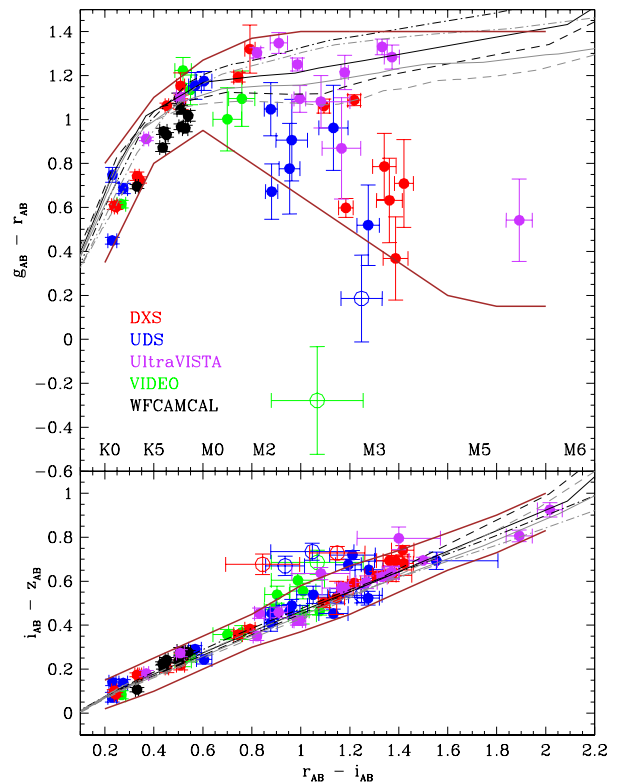


Figure 2. Colours of the likely stellar sources. No reddening correction has been applied. Open circles are objects with one or more colours that deviate from the loci identified by the sample. The thinner lines are model isochrones from the BT-SETTL set (black, Allard et al. (2012); Baraffe et al. (2015)) and the PARSEC set (grey, Bressan et al. (2012); Girardi et al. (2000)). For these sequences, age in Gyr and $[m/H]$ are: 5, 0.0 (solid line); 1.0, +0.3 (dashed line); 10, -0.5 (dash-dot line). The thick brown lines indicate the regions adopted here as defining normal star colours. Two sources are excluded in the top panel, being either too blue in $g_{AB} - r_{AB}$ or in $r_{AB} - i_{AB}$. One of these is also an outlier in the lower panel, along with four other sources that appear too blue in $r_{AB} - i_{AB}$ or too red in $i_{AB} - z_{AB}$. Spectral types along the x axis are from the relationship between $r_{AB} - i_{AB}$ and type given by Covey et al. (2007).

these are listed in Appendix Table C2. All of these rejected sources are at the faint end of the associated survey.

After exclusion of these 11 objects, the final sample consists of 81 stars ranging in spectral type from K2 to M6.

8 FINAL SAMPLE

Table 2 compiles the Pan-STARRS, UKIDSS/VISTA and Gaia data for the final sample of 81 stars. Spectral types are also given; these are estimated from the $r_{AB} - i_{AB}$ color (Covey et al. 2007) where available. Where this color was not available the type was estimated by interpolating the $i_{AB} - z_{AB}$ or $y_{AB} - J_{Vega}$ colors, using stars with all three colors to define the relationships.

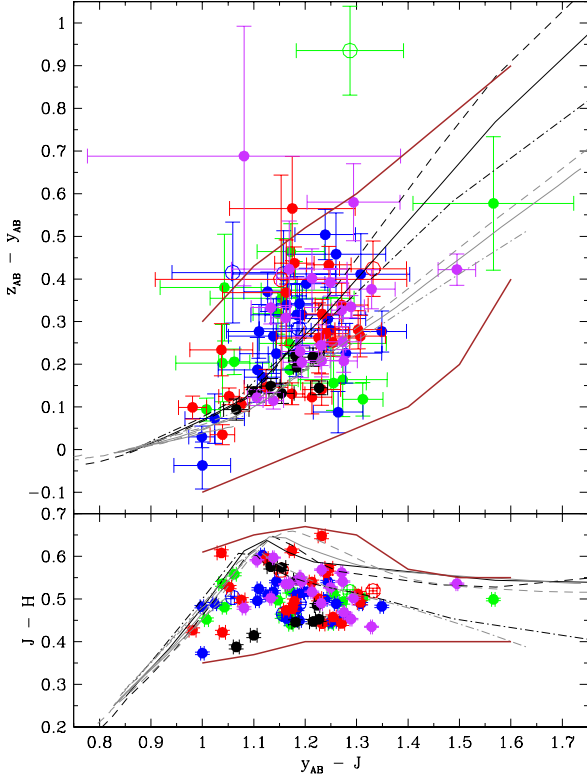


Figure 3. Colours of the likely stellar sources with symbols and lines as in Figure 2. No reddening correction has been applied. In addition to the 6 outliers identified in Figure 2, one additional outlier is identified in the upper panel, as too red in $z_{AB} - y_{AB}$ or too blue in $y_{AB} - J$.

Table 2. Data for Final Sample

IR Survey Name	RA ^o Decl. ^o	Spectral Type	RA ^o ± mas	Decl. ^o ± mas	Gaia Parallax ± mas	μ RA ± mas yr ⁻¹	μ Decl. ± mas yr ⁻¹	G	Pan-STARRS meanPSF AB mags					Z	VISTA Vega mags					
									<i>g</i> ±	<i>r</i> ±	<i>i</i> ±	<i>z</i> ±	<i>y</i> ±		<i>Y</i> ±	<i>J</i> ±	<i>H</i> ±	<i>Ks</i> ±		
UDS	34.0342392 -5.1728814	M3.5								21.7683	20.5566	19.8377	19.5578			18.3094	17.8626	17.5573		
UDS	34.0585392 -5.3530358	M3								0.0918	0.0133	0.0158	0.0340			0.0020	0.0018	0.0032		
UDS	34.0906904 -5.1871909	M2.5	34.090724762816 1.72235	-5.187250050454 2.62332				20.8948	22.0173	21.1110	20.1476	19.6591	19.5706			18.3070	17.8586	17.5862		
UDS	34.1166206 -5.4330768	M0	34.116703956194 1.13832	-5.433148535694 1.14117		31.4547	-14.0542	20.5484	21.6757	20.5011	19.8962	19.6520	19.4654			18.3581	17.8624	17.6619		
UDS	34.1459342 -5.4192513	M3				1.9471	1.7836		0.0312	0.0299	0.0088	0.0175	0.0736			0.0020	0.0018	0.0032		
UDS	34.1559576 -5.1864870	M3							21.9087	20.6313	19.9788	19.5901			18.3878	17.8711	17.5971			
UDS	34.1991747 -4.9476391	M2	34.199178767942 2.11700	-4.947697715516 2.55827				20.8574	22.0767	21.0300	20.1533	19.7429	19.4767			0.0021	0.0020	0.0033		
UDS	34.2232606 -4.7491396	M2.5							22.0736	21.2984	20.3444	19.8807	19.6561			18.5122	17.9708	17.7333		
UDS	34.2308454 -5.3911233	M3							0.2023	0.0397	0.0146	0.0271	0.0661			0.0022	0.0021	0.0034		
UDS	34.3103343 -5.4172864	M4.5							21.7580	20.5639	19.8892	19.5468			18.3580	17.9076	17.6218			
UDS	34.3240632 -5.4334966	M5							0.0466	0.0234	0.0114	0.0442			0.0021	0.0020	0.0032			
UDS	34.3475622 -5.4844981	M3						22.1287	21.6100	20.3351	19.8115	19.5857			18.5593	18.0758	17.7539			
UDS	34.3750969 -5.4848715	M4						0.1780	0.0413	0.0208	0.0229	0.0656			18.3072	17.8462	17.5767			
UDS	34.4389254 -5.4868910	M3							22.2721	20.7193	20.0259	19.7191			0.0020	0.0018	0.0032			
UDS	34.5319836 -5.1015107	M2							0.2496	0.0288	0.0261	0.0687			18.4748	17.9338	17.6445			
UDS	34.6700469 -5.0787756	M5.5							21.5469	20.3938	19.8619	19.4921			18.3653	17.8549	17.6085			
UDS	34.6776197 -4.9671258	M2.5							0.0350	0.0244	0.0254	0.0394			0.0021	0.0020	0.0033			
UDS	34.7079081 -4.9039390	K3	34.707936489500 1.49577	-4.903956579988 1.02274		6.0938	-0.6828	19.8705	21.7700	21.0978	20.2171	19.7631	19.4859			18.3762	17.8516	17.6315		
UDS	34.7425103 -5.4781043	K2	34.742551524244 0.84654	-5.478149266917 0.62984		-1.6880	-2.0504	19.7838	0.1245	0.0211	0.0142	0.0218	0.0591			0.0020	0.0018	0.0033		
UDS	34.7745424 -4.8889668	M5				1.3900	1.0100		20.8852	20.0903	19.5863				18.3469	17.8485	17.5109			
UDS	34.8172698 -4.7942724	M4							0.0378	0.0342	0.0477				0.0024	0.0020	0.0036			
UDS	34.8182276 -5.1062739	K7	34.818219802904 1.77157	-5.106312522351 1.21047				20.6559	21.5419	20.4913	19.9533	19.6368			18.4426	17.9281	17.6586			
UDS	34.8710949 -5.1154861	K2	34.871103834372 0.67365	-5.115517311973 0.56550		6.0251	-7.9278	19.6855	0.0643	0.0223	0.0325	0.0645			0.0022	0.0019	0.0035			
VIDEO	35.0542019 -4.4813805	M0				0.9709	0.8974		19.6602	19.3849	19.2484	19.1746			18.1506	17.6618	17.5555			
VIDEO	35.1152774 -4.9003215	M4.5							0.0188	0.0144	0.0105	0.0102	0.0573			0.0017	0.0017	0.0029		
VIDEO	35.3795003 -4.6057616	M2.5							20.4217	19.6735	19.4427	19.3034	19.2729			18.2738	17.7916	17.6957		
VIDEO	35.6598904 -4.6351333	K7	35.659947796971 0.97484	-4.635110377207 1.29221	2.9760	-1.5893	-8.0207	20.4536	0.0316	0.0109	0.0086	0.0109	0.0226			0.0019	0.0018	0.0029		
VIDEO	35.6610332 -4.5453215	M2.5							20.9764	20.1684	19.7567				18.4493	17.9604	17.6400			
VIDEO	35.6959814 -4.5264450	K7							0.0321	0.0280	0.0913				0.0022	0.0020	0.0034			
VIDEO	35.7281547 -5.2037306	M3.5							20.8266	20.1456	19.8074				18.6426	18.1886	17.9029			
VIDEO	35.8660837 -4.1075563	M3							0.0220	0.0338	0.0365				0.0022	0.0019	0.0033			
VIDEO	35.9623324 -4.2237615	M2							20.5580	19.9542	19.7055				18.5033	17.9004	17.7308			
VIDEO	35.9681759 -4.2573895	M2.5							0.0558	0.0243	0.0099	0.0119	0.0328		0.0021	0.0019	0.0034			
VIDEO	35.9729949 -5.2047350	M1	35.972994037179 2.08523	-5.204738379604 1.37865		5.0723	2.6179	20.7741	20.0623	19.6130	19.3828	19.3110	19.3479		18.3481	17.9747	17.8797			
VIDEO	36.0194131 -4.8728386	M6							0.0075	0.0131	0.0133	0.0175	0.0523		0.0016	0.0019	0.0029			
VIDEO	36.7075066 -4.5302703	K2	36.707528955400 0.79715	-4.530284911233 0.72506		-0.4147	-9.8394	20.0023	21.9615	20.9614	20.2612	19.9043	19.7169		19.0018	18.5465	18.0267	17.8310		
CAL	87.7352518 15.8816295	K3	87.735239861645 0.15854	15.881631420747 0.14113		-0.8779	-2.5404	18.2273	0.1321	0.0562	0.0170	0.0159	0.0324		0.0021	0.0028	0.0047	0.0043		
CAL	88.0439399 16.2125856	K5	88.043961153500 0.59256	16.212597884799 0.52929	0.3641	2.3316	-5.6403	18.7329	20.8536	20.1628	19.6987				19.0704	18.5271	18.0857	17.8255		
UVISTA	149.7053794 2.4862294	M3	149.705377978507 6.20199	2.486235956106 11.01370				20.7270	0.0245	0.0260	0.0604				0.0020	0.0021	0.0032	0.0049		
UVISTA	149.7087571 2.7888722	M2.5	149.708761363556 3.41061	2.78887908216031 5.09192				20.6709	21.5975	20.5879	20.0334	19.6534			19.1025	18.6101	18.1286	17.9108		
									0.0449	0.0275	0.0215	0.1227			0.0023	0.0023	0.0028	0.0049		
									20.4112	19.8614	19.5991	19.3927			18.7513	18.3307	17.7733	17.6104		
									0.0663	0.0149	0.0130	0.0262	0.0153		0.0016	0.0020	0.0030	0.0032		
									21.5461	20.5580	19.9542	19.7055			19.0260	18.5364	18.0755	17.8501		
									0.1076	0.0137	0.0262	0.0518			0.0019	0.0021	0.0025	0.0039		
									22.1388	20.9156	20.3969	20.1589	19.9562		19.3279	18.9173	18.3808	18.1880		
									0.0521	0.0287	0.0167	0.0209	0.0883		0.0026	0.0036	0.0043	0.0049		
											21.0309	20.3774	20.0223		19.3492	18.8672	18.4067	18.1372		
											0.0243	0.0327	0.1362		0.0027	0.0029	0.0041	0.0053		
									21.5821	20.5765	20.1285	19.8094			19.1724	18.6624	18.1648	17.9286		
									0.0493	0.0305	0.0296	0.0382			0.0024	0.0024	0.0042	0.0050		
									21.1034	20.2167	19.7468	19.6287			18.8035	18.3172	17.8155	17.5798		
									0.0327	0.0131	0.0196	0.0342			0.0017	0.0021	0.0034	0.0045		
									21.4215	20.3440	19.8795	19.6303			18.8680	18.3799	17.9367	17.7016		
									0.0414	0.0137	0.0135	0.1066			0.0021	0.0028	0.0028	0.0042		
									18.7733	18.2901	17.7774	17.5529			18.7733	18.2901	17.7774	17.5529		
									0.002167											

Table 2 – *continued* Data for Final Sample

IR Survey Name	Spectral Type	RA ^o Decl. ^o	RA ^o ± mas	Gaia Decl. ^o Parallax ± mas	μ RA ± mas yr ⁻¹	μ Decl. ± mas yr ⁻¹	G	Pan-STARRS meanPSF AB mags	VISTA Vega mags	<i>Ks</i>
								<i>g</i> <i>r</i> <i>i</i> <i>z</i> <i>y</i>	<i>Z</i> <i>Y</i> <i>J</i> <i>H</i> <i>Ks</i>	
UVISTA 149.7234241 2.7511805	M3.5							22.7104 21.3098 20.5154 19.9353	19.2576 18.6410 18.1390 17.8376	
UVISTA 149.7264829 2.5902461	M3.5	149.726469124665	2.590250565341				20.8861 22.7546 21.4708 20.0987 19.4587 19.0669	0.1649 0.0422 0.0297 0.0848	0.0012 0.0012 0.0019 0.0021	
UVISTA 149.7307359 2.7399894	M3.5							0.0478 0.0270 0.0148 0.0238 0.0208	18.3791 17.8180 17.2913 17.0327	
UVISTA 149.7313560 2.6928870	M2	149.731351713864	2.692886882358	2.2849	-5.1377	-2.4384	19.9948 21.4538 20.1508 19.3297 18.9789 18.7493	21.8113 20.4456 19.8085 19.4070	18.7503 18.1943 17.6766 17.4237	
UVISTA 149.731965 2.7327131	M4	0.63176	0.62314	0.8261	1.2600	1.1541		0.0307 0.0130 0.0246 0.0639	0.0010 0.0008 0.0013 0.0019	
UVISTA 149.7332239 2.7947267	M5.5							21.7754 21.0125 20.3254	19.8434 19.2441 18.7651 18.4652	
UVISTA 149.7355822 2.3205009	M2.5							0.0568 0.0570 0.2994	0.0022 0.0020 0.0025 0.0033	
UVISTA 149.7360512 2.4872598	K4	149.736045993165	2.487259844103	-0.2321	-7.3325	-5.4547	19.5169 20.3606 19.4492 19.0804 18.8997 18.7779	22.8180 22.2761 20.3844 19.5777 19.2018	18.4389 17.8730 17.4382 17.1655	
UVISTA 149.7476065 2.7272746	M2.5	0.64011	0.37359	0.9176	0.8382	0.8325		0.1809 0.0487 0.0215 0.0164 0.0430	0.0009 0.0008 0.0013 0.0016	
UVISTA 149.7564883 2.7279976	M3.5	149.756487939153	2.727999832329					22.3015 21.2084 20.2115 19.7921 19.5830	18.8246 18.3514 17.7820 17.5625	
UVISTA 149.7597522 2.7827541	M6	2.92877	5.22498					0.0571 0.0209 0.0122 0.0141 0.0365	0.0010 0.0007 0.0012 0.0014	
UVISTA 149.7742789 2.6687115	K7	149.774279147971	2.668706544853	-0.3755	-6.9845	-12.8963	19.8543 20.9477 19.8446 19.3370 19.0638 18.9487	0.0066 0.0042 0.0089 0.0113 0.0148	18.0802 17.6720 17.0809 16.9594	
UVISTA 149.7772394 2.7783479	M3	0.63748	0.58222	0.7819	1.0446	0.9791		22.9265 21.8455 20.7642 20.1292 19.8838	0.0008 0.0006 0.0009 0.0010	
UVISTA 149.7923194 2.5832843	M4	149.792307399543	2.583274196954					0.0222 0.1171 0.0209 0.0178 0.0443	19.1804 18.6534 18.1630 17.9131	
UVISTA 149.8156044 2.7697500	M2	149.815608819914	2.769757311106					20.5141 22.3374 21.0074 19.6753 19.0564 18.7287	0.0016 0.0014 0.0017 0.0027	
UVISTA 149.81712140 2.7717426	M2	149.817198369447	2.771754406963					0.0312 0.0179 0.0104 0.0098 0.0172	17.9949 17.4583 16.8964 16.6443	
UVISTA 149.8251519 2.7608222	M3	149.825158826105	2.760822268756					22.6823 20.6647 19.7397 19.3184	0.0011 0.0008 0.0010 0.0012	
CAL 276.7225557 4.0587376	K7	276.722547796086	4.058718034922	0.7440	1.2330	-1.1142	18.8017 19.9119 18.8163 18.3138 18.0759 17.8826 17.4972 17.2159	0.0444 0.0250 0.0208 0.0301	18.4764 17.8232 17.2872 16.9664	
CAL 276.8280740 3.9774041	K5	276.828044639757	3.977392480520	0.2437	0.4722	0.5167		0.0565 0.0098 0.0096 0.0238	0.0012 0.0009 0.0014 0.0015	
CAL 277.2399424 4.0952285	K7	277.239920836661	4.095191453128	0.4415	-1.0556	-4.3463	18.4684 19.3584 18.4862 18.0496 17.8239 17.6885 17.2663 17.0325	0.0131 0.0119 0.0033 0.0070 0.0089 0.0074 0.0035	18.2690 17.7772 17.2745 17.0546	
CAL 277.2699877 4.5221425	K7	277.269972705599	4.522125403396	0.1767	0.3142	0.3214		0.0351 0.0316 0.0148 0.0193 0.0332	0.0009 0.0010 0.0010 0.0013	
DXS 333.0711763 0.8303505	M1	333.071203143918	0.830304033314	0.3592	-4.1032	-2.9879	18.9311 19.9941 18.9491 18.4395 18.1591 18.0144 17.6120 17.3401	0.0587 0.0162 0.0122 0.0203 0.0301	18.2823 17.8016 17.2609 17.0372	
DXS 333.0715557 0.9162355	M3	333.071561677249	0.916173667634	0.3066	0.5879	1.1778		20.7489 22.2857 21.0721 19.8930 19.3248 19.1166	0.0009 0.0008 0.0013 0.0014	
DXS 333.0718253 1.1889522	K7	333.071813320305	1.188906051043	0.1228	21.2977	-2.2154	19.6927 20.9885 19.7961 19.0524 18.6951 18.5724	0.0777 0.0076 0.0134 0.0105 0.0236	17.3587 16.8191 16.5811	
DXS 333.2839533 -0.4105822	M3	333.283923388425	-0.410674164956	0.5681	1.0132	20.9872		0.0180 0.0141 0.0108 0.0117 0.0366	0.0027 0.0029 0.0046	
DXS 333.2928434 0.7343005	M3	333.292847124344	0.734269316808	1.2559	-0.0237	-4.0006	20.2384 21.3538 20.2010 19.6916 19.4740 19.2395	0.0362 0.0237 0.0173 0.0216 0.0790	18.1551 17.6539 17.3931	
DXS 333.3028790 0.7544730	M4	333.302887172419	0.754442344733	1.9038	1.6055	1.9193		0.0572 0.0135 0.0086 0.0194 0.0583	0.0035 0.0036 0.0054	
DXS 333.3594622 0.7512320	M3							21.3054 19.9960 19.3718 19.0911	17.7879 17.2757 17.0097	
DXS 333.3824122 0.7424491	M4							0.0441 0.0187 0.0091 0.0436	0.0030 0.0050 0.0048	
DXS 333.4515020 0.7278254	M3	333.451575578969	0.727800712380	-0.7282	20.6346	3.6240	20.6231 21.9420 21.1563 19.8150 19.1856 18.9350	0.0122 0.0223 0.0140 0.0034 0.0338	17.4070 16.9648 16.6916	
DXS 333.4987510 0.7349581	M4	0.86841	1.12594	1.3384	2.9172	2.2283		0.1965 0.0374 0.0103 0.0277 0.0427	0.0028 0.0033 0.0048	
DXS 333.7094552 0.5245924	K5	333.709446334135	0.524569397526	0.3710	-5.8741	-3.8524	19.2601 20.2572 19.1939 18.7410 18.5351 18.3885	21.5006 20.3548 19.8311 19.4634	17.5557 17.1132 16.8090	
DXS 333.9366041 -0.9471891	M1	333.936644901728	-0.947205672205	1.9967	0.3402	0.9006		0.0285 0.0143 0.0738 0.0992	0.0035 0.0037 0.0054	
DXS 334.6493784 1.3819891	M4							21.7048 20.2887 19.5484 18.9835	18.3002 17.8259 17.5602	
DXS 334.8182190 0.3928450	M4	334.818255423593	0.392802658682					0.0436 0.0132 0.0174 0.1214	0.0034 0.0034 0.0057	
DXS 334.9251383 0.7274321	K3	334.925138353308	0.727415208209	-0.8993	-0.8993	-5.5026	19.1814 19.8406 19.0993 18.7683 18.5935 18.4693	0.1427 0.0470 0.0100 0.0124 0.0300	17.6807 17.2226 16.9545	
DXS 335.0303119 0.4454373	K2	335.030321608241	0.445380424500	0.6877	0.7133	0.6893		21.9939 21.6259 20.2389 19.5468 19.1100	0.0044 0.0032 0.0049	
DXS 335.4746997 0.5091841	M2.5	335.474729551077	0.509174939517	0.40323	0.4482	0.1835		0.1835 0.0487 0.0139 0.0187 0.0334	17.9313 17.4356 17.1440	
DXS 335.5234567 0.2874459	K3	335.523439344855	0.287394859518	0.3402	0.3402	0.7032		0.0115 0.0043 0.0057 0.0215 0.0366	0.0034 0.0043 0.0052	
DXS 335.5823067 -0.1444080	K7	335.582301746661	-0.144465229042	1.0896	1.4193	-7.7910	19.7115 21.1755 19.8556 19.0643 18.6820 18.5408	0.0076 0.0056 0.0032 0.0067 0.0260	17.2650 16.6699 16.5297	
DXS 335.6150370 0.5091194	K2	335.615033009253	0.509092433091					21.5236 20.1513 19.5239 19.0910	0.0024 0.0058 0.0050	
		0.39047	0.28177					0.0793 0.0153 0.0244 0.0358	0.0039 0.0054 0.0051	
								20.7669 21.9217 21.2902 19.9285 19.2348 18.9691	17.8448 17.2731 16.9818	
								0.1852 0.0516 0.0094 0.0230 0.0427	0.0039 0.0054 0.0051	
								0.0234 0.0223 0.0070 0.0201 0.0375	17.6622 17.1697 16.8747	
								0.0234 0.0223 0.0070 0.0201 0.0375	0.0034 0.0044 0.0054	
								0.0234 0.0223 0.0070 0.0201 0.0375	17.4172 16.8889 16.7754	
								0.0234 0.0223 0.0070 0.0201 0.0375	0.0030 0.0016 0.0047	
								0.0234 0.0223 0.0070 0.0201 0.0375	0.0030 0.0016 0.0047	
								0.0234 0.0223 0.0070 0.0201 0.0375	17.4399 17.0128 16.9227	
								0.0234 0.0223 0.0070 0.0201 0.0375	0.0040 0.0035 0.0049	
								0.0234 0.0223 0.0070 0.0201 0.0375	17.3604 16.7993 16.5495	
								0.0234 0.0223 0.0070 0.0201 0.0375	0.0031 0.0034 0.0049	
								0.0234 0.0223 0.0070 0.0201 0.0375	17.1721 16.6742 16.5423	
								0.0234 0.0223 0.0070 0.0201 0.0375	0.0036 0.0036 0.0052	
								0.0234 0.0223 0.0070 0.0201 0.0375	17.3728 16.7596 16.6027	
								0.0234 0.0223 0.0070 0.0201 0.0375	0.0031 0.0034 0.0042	
								0.0234 0.0223 0.0070 0.0201 0.0375	17.2914 16.8690 16.7655	
								0.0234 0.0223 0.0070 0.0201 0.0375	0.0027 0.0036 0.0054	

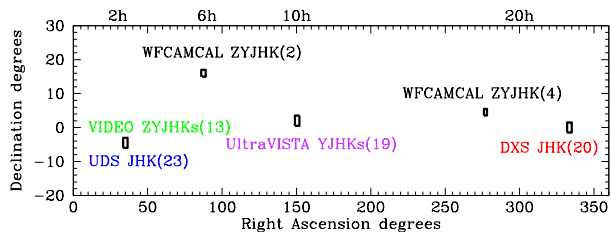


Figure 4. Sky chart showing the location and filter coverage of the final selection of standards. The number of stars at each location is given in parentheses.

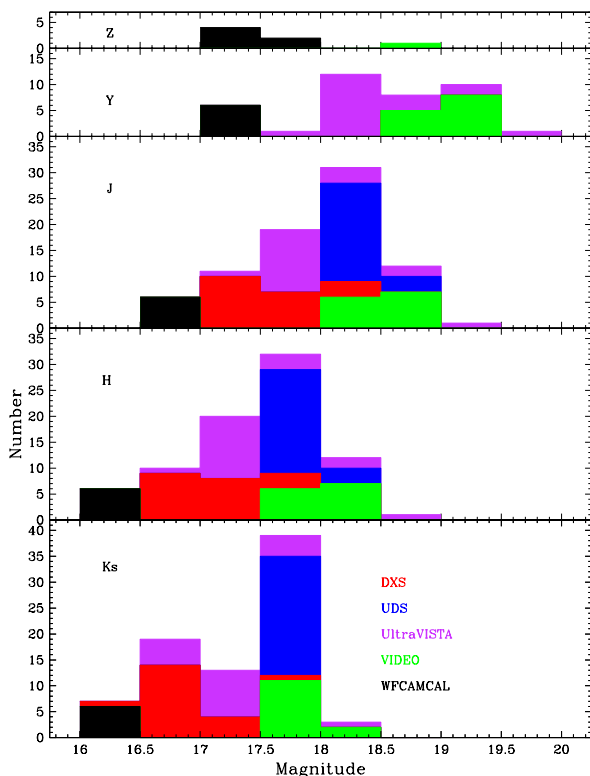


Figure 5. Histogram of magnitude distribution in each filter for the final selection of standard stars.

9 CALIBRATION DATA

Tables 3 and 4 list the photometric calibration data for Z and Y, and Tables 5 and 6 list the photometric calibration data for *JHK(s)*, selected from the WFCAM (UKIDSS) and VISTA archives respectively. The transformations between the WFCAM and VISTA systems are given in Section 4. The uncertainties in the magnitudes which have been converted to the VISTA system includes the uncertainty in the transformation.

We only include stars with twenty or more good observations in the variability tables, for each filter, for use as calibrators. Figure 4 illustrates the location on the sky of the new photometric standards, and Figure 5 shows the magnitude distribution for each filter.

The mean magnitude from the variability table, with an uncertainty given by $1.484 \times \text{MagMAD} / \sqrt{n\text{GoodObs}} - 1$, is the fundamental photometric reference for calibration purposes. The photometry presented here has a relative uncer-

tainty of < 0.006 mag. González-Fernández et al. (2018) and Hodgkin et al. (2009) compare the WFCAM and UKIDSS photometric systems, compare each system to the 2MASS system, and compare the WFCAM system to the FS system which is defined by a different camera on UKIRT. The authors also examine the VISTA and WFCAM colours of samples of A0 stars, which by definition are zero. These various comparisons suggest that the VISTA and WFCAM photometric accuracy is $\lesssim 0.02$ mag, which also applies to the sample of stars presented here.

10 CONCLUSIONS

The traditional near-infrared photometric standard stars published by Hawarden et al. (2001) and Leggett et al. (2006) are too bright for efficient observing by current 8- to 10-m class telescopes, and future extremely large 30- to 40-m class telescopes. In order to provide a set of fainter standards for community use, we queried the DXS, UDS, WFCAMCAL, VIDEO and UltraVISTA surveys using the WFCAM and VISTA Science Archives. Non-variable, probable stars were identified with *JHK(s)* magnitudes between 16 and 19. The initial sample of 169 sources was further refined by excluding 77 objects that may be galaxies, as indicated by red/near-infrared colours or PAN-STARRS Kron magnitudes. Eleven of the remaining sources were excluded due to atypical colours, or discrepant aperture and mean magnitudes from the source and variability tables, respectively. In this way we have produced a sample of 81 non-variable objects with precise photometry, that are likely to be single K and M stars. The new standard stars are distributed equatorially and are accessible from both hemispheres. Table 7 collates the calibration data on the VISTA *ZYJHKs* photometric system. The table also gives a unique running identification number for convenience — we refer to this as the Very Faint Standard, or VFS, identification number. We also give standard IAU names based on sexagesimal coordinate strings. Finder charts are presented in Appendix D.

Table 3. Standard Stars from WFCAM: ZY

Survey	RA [°] Decl. [°]	n Z	Z mean ± mag	Z Aper ± mag	Z rms mag	Z MAD mag	n Y	Y mean ± mag	Y Aper ± mag	Y rms mag	Y MAD mag	ZVISTA ± mag	YVISTA ± mag
CAL	87.7352518 15.8816295	106	17.2521 0.0026	17.2224 0.0149	0.0732	0.0177	121	17.0335 0.0025	17.0391 0.0138	0.0339	0.0183	17.2730 0.0070	17.0544 0.0035
CAL	88.0439399 16.2125856	86	17.4721 0.0053	17.4935 0.0178	0.1718	0.0327	103	17.2724 0.0032	17.2371 0.0154	0.0411	0.0218	17.4837 0.0095	17.2840 0.0044
CAL	276.7225557 4.0587376	133	17.4837 0.0028	17.4682 0.0173	0.0588	0.0218	131	17.2023 0.0030	17.1913 0.0151	0.0544	0.0227	17.4972 0.0081	17.2159 0.0042
CAL	276.8280740 3.9774041	140	17.2483 0.0026	17.2810 0.0155	0.0845	0.0203	138	17.0144 0.0023	17.0237 0.0137	0.0425	0.0181	17.2663 0.0074	17.0325 0.0035
CAL	277.2399424 4.0952285	135	17.5592 0.0038	17.5412 0.0182	0.0638	0.0294	137	17.2465 0.0028	17.2558 0.0160	0.0497	0.0221	17.5724 0.0085	17.2598 0.0041
CAL	277.2699877 4.5221425	135	17.5972 0.0033	17.5926 0.0183	0.0667	0.0257	133	17.3253 0.0028	17.2961 0.0157	0.0560	0.0218	17.6120 0.0081	17.3401 0.0040

Table 4. Standard Stars from VISTA: ZY

Survey	RA [°] Decl. [°]	n Z	Z mean ± mag	Z Aper ± mag	Z rms mag	Z MAD mag	n Y	Y mean ± mag	Y Aper ± mag	Y rms mag	Y MAD mag
VIDEO	35.0542019 -4.4813805						78	19.0018 0.0021	18.9692 0.0098	0.0327	0.0126
VIDEO	35.1152774 -4.9003215						78	19.0704 0.0020	19.0412 0.0088	0.0219	0.0118
VIDEO	35.3795003 -4.6057616						78	19.1025 0.0023	19.0890 0.0089	0.0294	0.0133
VIDEO	35.6598904 -4.6351333						77	18.7513 0.0016	18.7303 0.0052	0.0292	0.0094
VIDEO	35.6610332 -4.5453215						78	19.0260 0.0019	18.9936 0.0075	0.0234	0.0111
VIDEO	35.6959814 -4.5264450						78	19.3279 0.0026	19.2979 0.0105	0.0228	0.0154
VIDEO	35.7281547 -5.2037306						78	19.3492 0.0027	19.3460 0.0112	0.0236	0.0160
VIDEO	35.8660837 -4.1075563						78	19.1724 0.0024	19.1566 0.0106	0.0343	0.0143
VIDEO	35.9623324 -4.2237615						78	18.8035 0.0017	18.7944 0.0088	0.0206	0.0099
VIDEO	35.9681759 -4.2573895						77	18.8680 0.0021	18.8437 0.0288	0.0210	0.0122
VIDEO	35.9729949 -5.2047350						78	18.7733 0.0025	18.8435 0.0290	0.0211	0.0147
VIDEO	36.0194131 -4.8728386						78	19.1804 0.0022	19.1611 0.0144	0.0243	0.0128
VIDEO	36.7075066 -4.5302703	74	18.9232 0.0017	18.8959 0.0072	0.0460	0.0099	75	18.7517 0.0016	18.7244 0.0074	0.0157	0.0092
UVISTA	149.7053794 2.4862294						90	18.3682 0.0013	18.3562 0.0013	0.0122	0.0084
UVISTA	149.7087571 2.7888722						75	18.4555 0.0011	18.4382 0.0020	0.0085	0.0063
UVISTA	149.7234241 2.7511805						75	19.2576 0.0012	19.2364 0.0029	0.0104	0.0071
UVISTA	149.7264829 2.5902461						140	18.3791 0.0013	18.3591 0.0014	0.0133	0.0101
UVISTA	149.7307359 2.7399894						75	18.7503 0.0010	18.7310 0.0023	0.0091	0.0058
UVISTA	149.7313560 2.6928870						143	18.0241 0.0008	18.0070 0.0011	0.0094	0.0066
UVISTA	149.7319650 2.7327131						75	19.8434 0.0022	19.8107 0.0034	0.0181	0.0126
UVISTA	149.7332239 2.7947267						71	18.4389 0.0009	18.4214 0.0020	0.0086	0.0053
UVISTA	149.7355822 2.3205009						140	18.8246 0.0010	18.8101 0.0017	0.0121	0.0077
UVISTA	149.7360512 2.4872598						147	18.0802 0.0008	18.0609 0.0012	0.0109	0.0069
UVISTA	149.7476065 2.7272746						75	19.1804 0.0016	19.1586 0.0025	0.0140	0.0093
UVISTA	149.7564883 2.7279976						75	17.9949 0.0011	17.9755 0.0014	0.0098	0.0063
UVISTA	149.7597522 2.7827541						75	18.4764 0.0012	18.4621 0.0021	0.0097	0.0069
UVISTA	149.7742789 2.6687115						79	18.2324 0.0011	18.2136 0.0013	0.0106	0.0064
UVISTA	149.7772394 2.7783479						75	18.8640 0.0011	18.8447 0.0025	0.0097	0.0065
UVISTA	149.7923194 2.5832843						140	18.1225 0.0012	18.0959 0.0012	0.0123	0.0095
UVISTA	149.8156044 2.7697500						75	18.2690 0.0009	18.2479 0.0018	0.0093	0.0052
UVISTA	149.8172140 2.7717426						75	18.2823 0.0009	18.2632 0.0018	0.0075	0.0054
UVISTA	149.8251519 2.7608222						75	18.3345 0.0010	18.3074 0.0019	0.0102	0.0060

Note: UltraVISTA Y-band source table aperture magnitudes are systematically brighter than the variability table mean magnitudes by ≈ 0.02 mag, due to differences in the aperture correction and zeropoints. All mean magnitudes given in this paper are self-consistent.

Table 5. Standard Stars from WFCAM: *JHK*

Survey	RA ^o Decl. ^o	n <i>J</i>	<i>J</i> mean ± mag	<i>J</i> Aper ± mag	<i>J</i> rms mag	<i>J</i> MAD mag	n <i>H</i>	<i>H</i> mean ± mag	<i>H</i> Aper ± mag	<i>H</i> rms mag	<i>H</i> MAD mag	n <i>K</i>	<i>K</i> mean ± mag	<i>K</i> Aper ± mag	<i>K</i> rms mag	<i>K</i> MAD mag	<i>J</i> VISTA ± mag	<i>H</i> VISTA ± mag	<i>K</i> sVISTA ± mag
UDS	34.0342392 -5.1728814	349	18.3349 0.0009	18.5313 0.0025	0.0177	0.0108	225	17.8661 0.0013	18.0749 0.0022	0.0224	0.0130	513	17.5663 0.0009	17.7429 0.0019	0.0267	0.0141	18.3094 0.0020	17.8626 0.0018	17.5573 0.0032
UDS	34.0585392 -5.3530358	350	18.5195 0.0009	18.7181 0.0028	0.0180	0.0108	220	18.0215 0.0016	18.2202 0.0024	0.0238	0.0158	522	17.7618 0.0011	17.9339 0.0021	0.0286	0.0173	18.4943 0.0020	18.0184 0.0020	17.7527 0.0032
UDS	34.0906904 -5.1871909	351	18.3316 0.0009	18.5312 0.0026	0.0201	0.0117	225	17.8608 0.0015	18.0636 0.0023	0.0253	0.0154	516	17.5957 0.0010	17.7785 0.0020	0.0273	0.0147	18.3070 0.0020	17.8586 0.0020	17.5862 0.0031
UDS	34.1166206 -5.4330768	350	18.3820 0.0009	18.5803 0.0027	0.0187	0.0119	221	17.8637 0.0014	18.0607 0.0023	0.0226	0.0137	522	17.6719 0.0012	17.8423 0.0020	0.0295	0.0191	18.3581 0.0020	17.8624 0.0018	17.6619 0.0032
UDS	34.1459342 -5.4192513	352	18.4145 0.0010	18.6036 0.0028	0.0196	0.0125	221	17.8761 0.0012	18.0775 0.0023	0.0231	0.0124	524	17.6053 0.0011	17.7783 0.0020	0.0264	0.0173	18.3878 0.0021	17.8711 0.0018	17.5971 0.0033
UDS	34.1559576 -5.1864870	352	18.5141 0.0010	18.7003 0.0029	0.0228	0.0129	224	17.9824 0.0015	18.1923 0.0025	0.0271	0.0152	516	17.7474 0.0013	17.9236 0.0022	0.0304	0.0199	18.4886 0.0021	17.9790 0.0020	17.7385 0.0033
UDS	34.1991747 -4.9476391	353	18.3649 0.0011	18.5502 0.0028	0.0199	0.0138	221	17.8259 0.0014	18.0177 0.0022	0.0248	0.0143	524	17.5907 0.0011	17.7792 0.0020	0.0254	0.0170	18.3392 0.0021	17.8223 0.0019	17.5819 0.0033
UDS	34.2232606 -4.7491396	386	18.5385 0.0010	18.7258 0.0028	0.0227	0.0139	257	17.9753 0.0017	18.1769 0.0023	0.0291	0.0184	604	17.7418 0.0013	17.9272 0.0022	0.0344	0.0215	18.5122 0.0022	17.9708 0.0021	17.7333 0.0034
UDS	34.2308454 -5.3911233	351	18.3830 0.0010	18.5733 0.0023	0.0206	0.0130	222	17.9104 0.0016	18.1110 0.0021	0.0315	0.0162	535	17.6311 0.0012	17.8114 0.0019	0.0264	0.0186	18.3580 0.0021	17.9076 0.0020	17.6218 0.0032
UDS	34.3103343 -5.4172864	348	18.5864 0.0011	18.7776 0.0030	0.0259	0.0143	227	18.0813 0.0013	18.2905 0.0026	0.0252	0.0132	533	17.7619 0.0011	17.9373 0.0021	0.0285	0.0169	18.5593 0.0022	18.0758 0.0018	17.7539 0.0034
UDS	34.3240632 -5.4334966	345	18.5039 0.0012	18.6955 0.0029	0.0211	0.0149	227	17.9932 0.0015	18.2027 0.0024	0.0273	0.0154	530	17.6775 0.0011	17.8517 0.0020	0.0298	0.0174	18.4768 0.0023	17.9876 0.0020	17.6696 0.0034
UDS	34.3475622 -5.4844981	345	18.3321 0.0009	18.5255 0.0026	0.0178	0.0110	227	17.8488 0.0015	18.0557 0.0022	0.0205	0.0128	531	17.5860 0.0010	17.7628 0.0019	0.0243	0.0150	18.3072 0.0020	17.8462 0.0018	17.5767 0.0032
UDS	34.3750969 -5.4848715	345	18.5027 0.0009	18.6924 0.0028	0.0183	0.0109	227	17.9403 0.0014	18.1475 0.0023	0.0221	0.0140	531	17.6520 0.0010	17.8285 0.0020	0.0240	0.0149	18.4748 0.0022	17.9338 0.0019	17.6445 0.0034
UDS	34.4389254 -5.4868910	339	18.3909 0.0011	18.5846 0.0027	0.0219	0.0131	222	17.8585 0.0015	18.0665 0.0022	0.0250	0.0155	517	17.6173 0.0013	17.8000 0.0020	0.0284	0.0192	18.3653 0.0021	17.8549 0.0020	17.6085 0.0033
UDS	34.5319836 -5.1015107	439	18.4015 0.0008	18.5871 0.0022	0.0195	0.0119	270	17.8547 0.0013	18.0507 0.0019	0.0263	0.0148	627	17.6405 0.0013	17.8118 0.0018	0.0321	0.0213	18.3762 0.0020	17.8516 0.0018	17.6315 0.0033
UDS	34.6700469 -5.0787756	374	18.3749 0.0013	18.5585 0.0024	0.0244	0.0165	247	17.8552 0.0015	18.0493 0.0020	0.0668	0.0160	576	17.5183 0.0016	17.7034 0.0019	0.0378	0.0265	18.3469 0.0024	17.8485 0.0020	17.5109 0.0036
UDS	34.6776197 -4.9671258	404	18.4691 0.0010	18.6542 0.0023	0.0231	0.0140	244	17.9327 0.0014	18.1342 0.0020	0.0330	0.0149	594	17.6670 0.0016	17.8461 0.0019	0.0447	0.0257	18.4426 0.0022	17.9281 0.0019	17.6586 0.0035
UDS	34.7079081 -4.9039390	347	18.1715 0.0008	18.3579 0.0022	0.0152	0.0095	229	17.6592 0.0012	17.8669 0.0018	0.0208	0.0123	533	17.5672 0.0011	17.7433 0.0018	0.0286	0.0166	18.1506 0.0017	17.6618 0.0017	17.5555 0.0029
UDS	34.7425103 -5.4781043	345	18.2942 0.0010	18.4874 0.0028	0.0200	0.0129	229	17.7883 0.0014	17.9945 0.0023	0.0249	0.0142	529	17.7077 0.0012	17.8798 0.0022	0.0293	0.0193	18.2738 0.0019	17.7916 0.0018	17.6957 0.0029
UDS	34.7745424 -4.8889668	371	18.4765 0.0010	18.6584 0.0024	0.0215	0.0125	260	17.9661 0.0015	18.1642 0.0021	0.0264	0.0162	595	17.6479 0.0013	17.8183 0.0019	0.0336	0.0210	18.4493 0.0022	17.9604 0.0020	17.6400 0.0034
UDS	34.8172698 -4.7942724	354	18.6678 0.0012	18.8510 0.0029	0.0221	0.0148	220	18.1916 0.0014	18.3840 0.0027	0.0243	0.0143	521	17.9121 0.0013	18.0983 0.0024	0.0347	0.0194	18.6426 0.0022	18.1886 0.0019	17.9029 0.0033
UDS	34.8182276 -5.1062739	404	18.5294 0.0009	18.7202 0.0025	0.0220	0.0122	249	17.9046 0.0014	18.1129 0.0020	0.0277	0.0150	619	17.7394 0.0013	17.9188 0.0020	0.0338	0.0221	18.5033 0.0021	17.9004 0.0019	17.7308 0.0034
UDS	34.8710949 -5.1154861	366	18.3653 0.0009	18.5585 0.0025	0.0178	0.0112	228	17.9672 0.0015	18.1734 0.0023	0.0259	0.0157	536	17.8937 0.0015	18.0786 0.0023	0.0387	0.0240	18.3481 0.0016	17.9747 0.0019	17.8797 0.0029
CAL	87.7352518 15.8816295	136	16.6289 0.0028	16.6552 0.0160	0.0591	0.0219	132	16.2164 0.0039	16.1917 0.0139	0.0522	0.0301	136	16.1131 0.0054	16.1481 0.0229	0.0639	0.0424	16.6104 0.0031	16.2224 0.0041	16.0999 0.0060
CAL	88.0439399 16.2125856	130	16.7948 0.0044	16.7686 0.0170	0.0543	0.0338	125	16.1988 0.0033	16.1922 0.0137	0.0559	0.0251	133	16.0276 0.0040	16.0285 0.0214	0.0613	0.0313	16.7693 0.0048	16.1954 0.0036	16.0186 0.0051
CAL	276.7225557 4.0587376	156	16.7231 0.0031	16.7228 0.0174	0.0678	0.0261	157	16.1566 0.0032	16.1678 0.0151	0.0612	0.0271	167	16.0093 0.0043	16.0034 0.0252	0.0665	0.0370	16.6992 0.0036	16.1551 0.0034	15.9994 0.0052
CAL	276.8280740 3.9774041	158	16.6085 0.0032	16.6102 0.0162	0.0485	0.0270	157	16.1697 0.0030	16.1597 0.0149	0.0612	0.0255	169	16.0157 0.0044	16.0139 0.0254	0.0750	0.0381	16.5879 0.0036	16.1728 0.0033	16.0038 0.0051
CAL	277.2399424 4.0952285	156	16.7699 0.0032	16.8010 0.0187	0.0727	0.0270	154	16.3020 0.0035	16.2767 0.0169	0.0543	0.0294	168	16.0474 0.0042	16.0938 0.0263	0.0716	0.0369	16.7456 0.0037	16.3002 0.0037	16.0376 0.0052
CAL	277.2699877 4.5221425	157	16.8103 0.0033	16.8261 0.0181	0.0741	0.0277	156	16.3346 0.0035	16.3425 0.0171	0.0621	0.0291	167	16.1284 0.0049	16.1660 0.0262	0.0715	0.0422	16.7872 0.0037	16.3344 0.0037	16.1180 0.0056

Note: UDS aperture magnitudes are not aperture corrected and are systematically fainter than the mean magnitudes by ≈ 0.2 mag.

Table 5 – continued Standard Stars from WFCAM: *JHK*

Survey	RA ^o Decl. ^o	n	<i>J</i> mean ± mag	<i>J</i> Aper ± mag	<i>J</i> rms mag	<i>J</i> MAD mag	n	<i>H</i> mean ± mag	<i>H</i> Aper ± mag	<i>H</i> rms mag	<i>H</i> MAD mag	n	<i>K</i> mean ± mag	<i>K</i> Aper ± mag	<i>K</i> rms mag	<i>K</i> MAD mag	<i>J</i> _{VISTA} ± mag	<i>H</i> _{VISTA} ± mag	<i>K</i> _{SVISTA} ± mag
DXS	333.0711763 0.8303505	26	17.3850 0.0020	17.3946 0.0071	0.0128	0.0067						22	16.5895 0.0034	16.5916 0.0068	0.0227	0.0105	17.3587 0.0027	16.8191 0.0029	16.5811 0.0046
DXS	333.0715557 0.9162355	26	18.1809 0.0029	18.1990 0.0106	0.0202	0.0099						22	17.4019 0.0045	17.4092 0.0115	0.0224	0.0138	18.1551 0.0035	17.6539 0.0036	17.3931 0.0054
DXS	333.0718253 1.1889522	26	18.2282 0.0037	18.2278 0.0109	0.0212	0.0125						22	17.4554 0.0050	17.4472 0.0129	0.0250	0.0155	18.2026 0.0042	17.5949 0.0055	17.4465 0.0059
DXS	333.2839533 -0.4105822	24	17.8141 0.0023	17.8167 0.0089	0.0133	0.0075						27	17.0182 0.0037	17.0135 0.0085	0.0258	0.0126	17.7879 0.0030	17.2757 0.0050	17.0097 0.0048
DXS	333.2928434 0.7343005	24	17.4314 0.0021	17.4294 0.0072	0.0159	0.0069	29	16.9668 0.0031	16.9427 0.0059	0.0145	0.0111	24	16.7012 0.0038	16.7006 0.0074	0.0245	0.0124	17.4070 0.0028	16.9648 0.0033	16.6916 0.0048
DXS	333.3028790 0.7544734	24	17.5811 0.0029	17.5772 0.0077	0.0156	0.0095	29	17.1164 0.0035	17.1209 0.0062	0.0175	0.0124	24	16.8181 0.0045	16.8164 0.0077	0.0194	0.0145	17.5557 0.0035	17.1132 0.0037	16.8090 0.0054
DXS	333.3594622 0.7512320	24	18.3253 0.0029	18.3266 0.0114	0.0170	0.0095	29	17.8289 0.0032	17.8372 0.0095	0.0277	0.0114	24	17.5693 0.0048	17.5466 0.0123	0.0278	0.0156	18.3002 0.0034	17.8259 0.0034	17.5602 0.0057
DXS	333.3824122 0.7424491	24	17.8349 0.0020	17.8492 0.0089	0.0114	0.0064	29	17.3366 0.0034	17.3456 0.0071	0.0197	0.0122	24	17.0058 0.0036	17.0084 0.0087	0.0220	0.0115	17.8077 0.0028	17.3309 0.0037	16.9979 0.0048
DXS	333.4515020 0.7278254	24	17.7055 0.0040	17.6906 0.0082	0.0147	0.0131	29	17.2250 0.0029	17.2250 0.0070	0.0161	0.0104	24	16.9639 0.0039	16.9546 0.0086	0.0193	0.0126	17.6807 0.0044	17.2226 0.0032	16.9545 0.0049
DXS	333.4987510 0.7349581	24	17.9579 0.0029	17.9683 0.0094	0.0163	0.0092	29	17.4404 0.0041	17.4481 0.0075	0.0223	0.0146	24	17.1523 0.0042	17.1722 0.0096	0.0199	0.0134	17.9313 0.0034	17.4356 0.0043	17.1440 0.0052
DXS	333.7094552 0.5245924	25	17.2900 0.0016	17.2783 0.0066	0.0111	0.0052	27	16.6727 0.0030	16.6579 0.0050	0.0205	0.0103	22	16.5390 0.0057	16.5553 0.0066	0.0227	0.0175	17.2650 0.0024	16.6699 0.0032	16.5297 0.0064
DXS	333.9366041 -0.9471891	24	17.3359 0.0014	17.3307 0.0070	0.0087	0.0045						24	16.5021 0.0039	16.5124 0.0063	0.0240	0.0126	17.3086 0.0024	16.6614 0.0058	16.4942 0.0050
DXS	334.6493784 1.3819891	33	17.8736 0.0033	17.8719 0.0087	0.0140	0.0128	20	17.2809 0.0053	17.2940 0.0077	0.0240	0.0154	44	16.9888 0.0038	16.9910 0.0086	0.0264	0.0169	17.8448 0.0039	17.2731 0.0054	16.9818 0.0051
DXS	334.8182190 0.3928450	32	17.6887 0.0029	17.6897 0.0087	0.0187	0.0107						34	16.8830 0.0043	16.8723 0.0083	0.0267	0.0168	17.6622 0.0034	17.1697 0.0044	16.8747 0.0054
DXS	334.9251383 0.7274321	31	17.4395 0.0025	17.4246 0.0074	0.0122	0.0093	21	16.8880 0.0011	16.8930 0.0061	0.0102	0.0033	34	16.7863 0.0038	16.7858 0.0078	0.0329	0.0148	17.4172 0.0030	16.8889 0.0016	16.7754 0.0047
DXS	335.0303119 0.4454373	29	17.4586 0.0038	17.4514 0.0076	0.0215	0.0135	22	17.0071 0.0033	17.0150 0.0067	0.0140	0.0103	30	16.9359 0.0041	16.9255 0.0086	0.0234	0.0150	17.4399 0.0040	17.0128 0.0035	16.9227 0.0049
DXS	335.4746997 0.5091841	45	17.3876 0.0025	17.3823 0.0072	0.0200	0.0111	33	16.8051 0.0031	16.8104 0.0054	0.0159	0.0119	46	16.5574 0.0038	16.5396 0.0067	0.0301	0.0170	17.3604 0.0031	16.7993 0.0034	16.5495 0.0049
DXS	335.5234567 0.2874459	31	17.1940 0.0032	17.1661 0.0066	0.0169	0.0117	26	16.6729 0.0034	16.6621 0.0051	0.0166	0.0116	32	16.5534 0.0044	16.5204 0.0068	0.0242	0.0165	17.1721 0.0036	16.6742 0.0036	16.5423 0.0052
DXS	335.5823067 -0.1444080	28	17.3988 0.0024	17.4093 0.0074	0.0155	0.0086						33	16.6114 0.0028	16.6016 0.0070	0.0201	0.0106	17.3728 0.0031	16.7596 0.0034	16.6027 0.0042
DXS	335.6150576 0.5091194	31	17.3103 0.0023	17.3109 0.0070	0.0124	0.0083	26	16.8637 0.0034	16.8537 0.0056	0.0144	0.0114	32	16.7785 0.0048	16.7719 0.0078	0.0269	0.0180	17.2914 0.0027	16.8690 0.0036	16.7655 0.0054

Table 6. Standard Stars from VISTA: *JHKs*

Survey	RA ^o Decl. ^o	n	<i>J</i> mean ± mag	<i>J</i> Aper ± mag	<i>J</i> rms mag	<i>J</i> MAD mag	n	<i>H</i> mean ± mag	<i>H</i> Aper ± mag	<i>H</i> rms mag	<i>H</i> MAD mag	n	<i>Ks</i> mean ± mag	<i>Ks</i> Aper ± mag	<i>Ks</i> rms mag	<i>Ks</i> MAD mag
VIDEO	35.0542019 -4.4813805	39	18.5465 0.0028	18.5352 0.0093	0.0178	0.0117	40	18.0267 0.0047	18.0154 0.0087	0.0331	0.0198	48	17.8310 0.0043	17.8200 0.0097	0.0394	0.0199
VIDEO	35.1152774 -4.9003215	39	18.5271 0.0021	18.5256 0.0081	0.0174	0.0085	40	18.0857 0.0032	18.0768 0.0080	0.0243	0.0134	48	17.8255 0.0049	17.8019 0.0088	0.0366	0.0225
VIDEO	35.3795003 -4.6057616	39	18.6101 0.0023	18.5996 0.0081	0.0138	0.0096	40	18.1286 0.0028	18.1186 0.0077	0.0174	0.0120	48	17.9108 0.0049	17.8928 0.0085	0.0369	0.0225
VIDEO	35.6598904 -4.635133	39	18.3307 0.0020	18.3223 0.0050	0.0343	0.0083	40	17.7733 0.0030	17.7600 0.0047	0.0237	0.0126	48	17.6104 0.0032	17.5946 0.0054	0.0293	0.0146
VIDEO	35.6610332 -4.5453215	39	18.5364 0.0021	18.5202 0.0070	0.0155	0.0087	40	18.0755 0.0025	18.0640 0.0070	0.0200	0.0104	48	17.8501 0.0039	17.8392 0.0076	0.0403	0.0182
VIDEO	35.6959814 -4.5264450	39	18.9173 0.0036	18.9030 0.0101	0.0226	0.0148	40	18.3808 0.0043	18.3622 0.0097	0.0287	0.0180	48	18.1880 0.0049	18.1696 0.0107	0.0400	0.0228
VIDEO	35.7281547 -5.2037306	39	18.8672 0.0029	18.8681 0.0101	0.0221	0.0121	40	18.4067 0.0041	18.4025 0.0098	0.0310	0.0172	48	18.1372 0.0053	18.1443 0.0105	0.0398	0.0244
VIDEO	35.8660837 -4.1075563	39	18.6624 0.0024	18.6644 0.0098	0.0170	0.0100	40	18.1648 0.0042	18.1651 0.0094	0.0240	0.0175	48	17.9286 0.0050	17.9385 0.0104	0.0383	0.0229
VIDEO	35.9623324 -4.2237615	39	18.3172 0.0021	18.3222 0.0080	0.0195	0.0089	40	17.8155 0.0034	17.8129 0.0078	0.0239	0.0142	48	17.5798 0.0045	17.5857 0.0086	0.0345	0.0208
VIDEO	35.9681759 -4.2573895	39	18.3799 0.0028	18.3917 0.0247	0.0163	0.0116	40	17.9367 0.0028	17.9951 0.0325	0.0207	0.0116	48	17.7016 0.0042	17.7283 0.0279	0.0354	0.0193
VIDEO	35.9729949 -5.2047350	39	18.2901 0.0022	18.3199 0.0583	0.0252	0.0092	40	17.7774 0.0028	17.7690 0.0551	0.0205	0.0120	48	17.5529 0.0039	17.6046 0.0430	0.0367	0.0182
VIDEO	36.0194131 -4.8728386	39	18.4466 0.0024	18.4416 0.0119	0.0169	0.0098	40	17.9480 0.0021	17.9333 0.0121	0.0218	0.0089	48	17.6283 0.0044	17.6006 0.0116	0.0367	0.0203
VIDEO	36.7075066 -4.5302703	38	18.4144 0.0022	18.3932 0.0073	0.0273	0.0091	40	17.9624 0.0034	17.9428 0.0075	0.0301	0.0142	50	17.8677 0.0040	17.8408 0.0082	0.0336	0.0189
UVISTA	149.7053794 2.4862294	131	17.8524 0.0009	17.8119 0.0012	0.0095	0.0066	232	17.3109 0.0010	17.3465 0.0012	0.0362	0.0105	199	17.0759 0.0010	17.0883 0.0012	0.0175	0.0093
UVISTA	149.7087571 2.7888722	80	17.9656 0.0008	17.9106 0.0019	0.0086	0.0047	123	17.4313 0.0012	17.4534 0.0018	0.0155	0.0092	101	17.2116 0.0018	17.2032 0.0019	0.0182	0.0121
UVISTA	149.7234241 2.7511805	80	18.6410 0.0012	18.5871 0.0027	0.0117	0.0071	123	18.1390 0.0019	18.1468 0.0023	0.0238	0.0139	101	17.8376 0.0021	17.8220 0.0024	0.0237	0.0143
UVISTA	149.7264829 2.5902461	150	17.8180 0.0007	17.7721 0.0012	0.0094	0.0057	236	17.2913 0.0011	17.3231 0.0012	0.0750	0.0110	203	17.0327 0.0012	17.0421 0.0012	0.0169	0.0115
UVISTA	149.7307359 2.7399894	80	18.1943 0.0008	18.1414 0.0021	0.0083	0.0045	123	17.6766 0.0013	17.7059 0.0020	0.0194	0.0100	101	17.4237 0.0019	17.4190 0.0020	0.0194	0.0129
UVISTA	149.7313560 2.6928870	158	17.5590 0.0006	17.5094 0.0011	0.0075	0.0047	252	17.0082 0.0009	17.0394 0.0011	0.0394	0.0100	206	16.7933 0.0009	16.8107 0.0011	0.0159	0.0084
UVISTA	149.7319650 2.7327131	80	19.2441 0.0020	19.1790 0.0031	0.0188	0.0118	123	18.7651 0.0025	18.7666 0.0027	0.0308	0.0186	101	18.4652 0.0033	18.4630 0.0031	0.0324	0.0225
UVISTA	149.7332239 2.7947267	78	17.8730 0.0008	17.8189 0.0018	0.0084	0.0048	120	17.4382 0.0013	17.4644 0.0018	0.0167	0.0095	97	17.1655 0.0016	17.1668 0.0019	0.0183	0.0106
UVISTA	149.7355822 2.3205009	150	18.3514 0.0007	18.3066 0.0015	0.0107	0.0055	236	17.7820 0.0012	17.7993 0.0015	0.0887	0.0123	203	17.5625 0.0014	17.5589 0.0015	0.0215	0.0130
UVISTA	149.7360512 2.4872598	152	17.6720 0.0006	17.6253 0.0011	0.0087	0.0051	232	17.0809 0.0009	17.1141 0.0011	0.0622	0.0093	199	16.9594 0.0010	16.9764 0.0012	0.0186	0.0092
UVISTA	149.7476065 2.7272746	80	18.6534 0.0014	18.6047 0.0023	0.0125	0.0087	123	18.1630 0.0017	18.1692 0.0021	0.0262	0.0129	101	17.9131 0.0027	17.9004 0.0022	0.0306	0.0182
UVISTA	149.7564883 2.7279976	80	17.4583 0.0008	17.4066 0.0013	0.0078	0.0048	123	16.8964 0.0010	16.9283 0.0012	0.0164	0.0073	101	16.6443 0.0012	16.6641 0.0013	0.0150	0.0078
UVISTA	149.7597522 2.7827541	80	17.8232 0.0009	17.7737 0.0018	0.0084	0.0053	123	17.2872 0.0014	17.3208 0.0017	0.0191	0.0106	101	16.9664 0.0015	16.9830 0.0018	0.0159	0.0100
UVISTA	149.7742789 2.6687115	104	17.8107 0.0008	17.7615 0.0013	0.0100	0.0056	251	17.2140 0.0010	17.2393 0.0012	0.0799	0.0111	206	17.0529 0.0010	17.0630 0.0013	0.0172	0.0093
UVISTA	149.7772394 2.7783479	80	18.3290 0.0011	18.2735 0.0023	0.0110	0.0069	123	17.7934 0.0014	17.8162 0.0021	0.0248	0.0105	101	17.5463 0.0018	17.5351 0.0021	0.0218	0.0121
UVISTA	149.7923194 2.5832843	150	17.5830 0.0006	17.5310 0.0011	0.0076	0.0046	237	17.1304 0.0011	17.1662 0.0011	0.0826	0.0114	203	16.8712 0.0010	16.8883 0.0011	0.0161	0.0096
UVISTA	149.8156044 2.7697500	80	17.7772 0.0010	17.7230 0.0018	0.0097	0.0057	123	17.2745 0.0010	17.3082 0.0017	0.0174	0.0074	101	17.0546 0.0013	17.0646 0.0018	0.0184	0.0090
UVISTA	149.8172140 2.7717426	80	17.8016 0.0008	17.7487 0.0018	0.0074	0.0049	123	17.2609 0.0013	17.2904 0.0017	0.0202	0.0098	101	17.0372 0.0014	17.0491 0.0018	0.0165	0.0096
UVISTA	149.8251519 2.7608222	80	17.8417 0.0007	17.7810 0.0018	0.0074	0.0041	123	17.3732 0.0012	17.3980 0.0017	0.0165	0.0089	101	17.1457 0.0014	17.1466 0.0018	0.0187	0.0097

Note: UltraVISTA source table aperture magnitudes are systematically brighter than the variability table mean magnitudes by ≈ 0.05 mag at *J*, and fainter by ≈ 0.02 mag at *H*, due to differences in the aperture correction and zeropoints. All mean magnitudes given in this paper are self-consistent.

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Table 7. Standard Stars with VISTA-System *ZYJHKs*. The first and second columns give a unique running number for convenience (VFS = ‘Very Faint Standard’) and a formal IAU designation for each star.

VFS	IAU name	RA [°] Decl. [°]	n Z	Z n Y ± mag	Y n J ± mag	J n H ± mag	H n Ks ± mag	Ks ± mag
1	UUDS J021608.22-051022.4	34.0342392			349 18.3094	225 17.8626	513 17.5573	
		-5.1728814			0.0020	0.0018	0.0032	
2	UUDS J021614.05-052110.9	34.0585392			350 18.4943	220 18.0184	522 17.7527	
		-5.3530358			0.0020	0.0020	0.0032	
3	UUDS J021621.77-051113.9	34.0906904			351 18.3070	225 17.8586	516 17.5862	
		-5.1871909			0.0020	0.0020	0.0031	
4	UUDS J021627.99-052559.1	34.1166206			350 18.3581	221 17.8624	522 17.6619	
		-5.4330768			0.0020	0.0018	0.0032	
5	UUDS J021635.02-052509.3	34.1459342			352 18.3878	221 17.8711	524 17.5971	
		-5.4192513			0.0021	0.0018	0.0033	
6	UUDS J021637.43-051111.4	34.1559576			352 18.4886	224 17.9790	516 17.7385	
		-5.1864870			0.0021	0.0020	0.0033	
7	UUDS J021647.80-045651.5	34.1991747			353 18.3392	221 17.8223	524 17.5819	
		-4.9476391			0.0021	0.0019	0.0033	
8	UUDS J021653.58-044456.9	34.2232606			386 18.5122	257 17.9708	604 17.7333	
		-4.7491396			0.0022	0.0021	0.0034	
9	UUDS J021655.40-052328.0	34.2308454			351 18.3580	222 17.9076	535 17.6218	
		-5.3911233			0.0021	0.0020	0.0032	
10	UUDS J021714.48-052502.2	34.3103343			348 18.5593	227 18.0758	533 17.7539	
		-5.4172864			0.0022	0.0018	0.0034	
11	UUDS J021717.78-052600.6	34.3240632			345 18.4768	227 17.9876	530 17.6696	
		-5.4334966			0.0023	0.0020	0.0034	
12	UUDS J021723.41-052904.2	34.3475622			345 18.3072	227 17.8462	531 17.5767	
		-5.4844981			0.0020	0.0018	0.0032	
13	UUDS J021730.02-052905.5	34.3750969			345 18.4748	227 17.9338	531 17.6445	
		-5.4848715			0.0022	0.0019	0.0034	
14	UUDS J021745.34-052912.8	34.4389254			339 18.3653	222 17.8549	517 17.6085	
		-5.4868910			0.0021	0.0020	0.0033	
15	UUDS J021807.68-050605.4	34.5319836			439 18.3762	270 17.8516	627 17.6315	
		-5.1015107			0.0020	0.0018	0.0033	
16	UUDS J021840.81-050443.6	34.6700469			374 18.3469	247 17.8485	576 17.5109	
		-5.0787756			0.0024	0.0020	0.0036	
17	UUDS J021842.63-045801.7	34.6776197			404 18.4426	244 17.9281	594 17.6586	
		-4.9671258			0.0022	0.0019	0.0035	
18	UUDS J021849.90-045414.2	34.7079081			347 18.1506	229 17.6618	533 17.5555	
		-4.9039390			0.0017	0.0017	0.0029	
19	UUDS J021858.20-052841.2	34.7425103			345 18.2738	229 17.7916	529 17.6957	
		-5.4781043			0.0019	0.0018	0.0029	
20	UUDS J021905.89-045320.3	34.7745424			371 18.4493	260 17.9604	595 17.6400	
		-4.8889668			0.0022	0.0020	0.0034	
21	UUDS J021916.14-044739.4	34.8172698			354 18.6426	220 18.1886	521 17.9029	
		-4.7942724			0.0022	0.0019	0.0033	
22	UUDS J021916.37-050622.6	34.8182276			404 18.5033	249 17.9004	619 17.7308	
		-5.1062739			0.0021	0.0019	0.0034	
23	UUDS J021929.06-050655.7	34.8710949			366 18.3481	228 17.9747	536 17.8797	
		-5.1154861			0.0016	0.0019	0.0029	
24	VIDEO J022013.01-042853.0	35.0542019		78 19.0018	39 18.5465	40 18.0267	48 17.8310	
		-4.4813805		0.0021	0.0028	0.0047	0.0043	
25	VIDEO J022027.67-045401.2	35.1152774		78 19.0704	39 18.5271	40 18.0857	48 17.8255	
		-4.9003215		0.0020	0.0021	0.0032	0.0049	
26	VIDEO J022131.08-043620.7	35.3795003		78 19.1025	39 18.6101	40 18.1286	48 17.9108	
		-4.6057616		0.0023	0.0023	0.0028	0.0049	
27	VIDEO J022238.37-043806.5	35.6598904		77 18.7513	39 18.3307	40 17.7733	48 17.6104	
		-4.6351333		0.0016	0.0020	0.0030	0.0032	
28	VIDEO J022238.65-043243.2	35.6610332		78 19.0260	39 18.5364	40 18.0755	48 17.8501	
		-4.5453215		0.0019	0.0021	0.0025	0.0039	
29	VIDEO J022247.04-043135.2	35.6959814		78 19.3279	39 18.9173	40 18.3808	48 18.1880	
		-4.5264450		0.0026	0.0036	0.0043	0.0049	
30	VIDEO J022254.76-051213.4	35.7281547		78 19.3492	39 18.8672	40 18.4067	48 18.1372	
		-5.2037306		0.0027	0.0029	0.0041	0.0053	
31	VIDEO J022327.86-040627.2	35.8660837		78 19.1724	39 18.6624	40 18.1648	48 17.9286	
		-4.1075563		0.0024	0.0024	0.0042	0.0050	
32	VIDEO J022350.96-041325.5	35.9623324		78 18.8035	39 18.3172	40 17.8155	48 17.5798	
		-4.2237615		0.0017	0.0021	0.0034	0.0045	
33	VIDEO J022352.36-041526.6	35.9681759		77 18.8680	39 18.3799	40 17.9367	48 17.7016	
		-4.2573895		0.0021	0.0028	0.0028	0.0042	
34	VIDEO J022353.52-051217.0	35.9729949		78 18.7733	39 18.2901	40 17.7774	48 17.5529	
		-5.2047350		0.0025	0.0022	0.0028	0.0039	
35	VIDEO J022404.66-045222.2	36.0194131		78 19.1804	39 18.4466	40 17.9480	48 17.6283	
		-4.8728386		0.0022	0.0024	0.0021	0.0044	
36	VIDEO J022649.80-043149.0	36.7075066	74 18.9232	75 18.7517	38 18.4144	40 17.9624	50 17.8677	
		-4.5302703	0.0017	0.0016	0.0022	0.0034	0.0040	
37	UCAL J055056.46+155253.9	87.7352518	106 17.2730	121 17.0544	136 16.6104	132 16.2224	136 16.0999	
		15.8816295	0.0070	0.0035	0.0031	0.0041	0.0060	
38	UCAL J055210.55+161245.3	88.0439399	86 17.4837	103 17.2840	130 16.7693	125 16.1954	133 16.0186	
		16.2125856	0.0095	0.0044	0.0048	0.0036	0.0051	
39	UltraVISTA J095849.29+022910.4	149.7053794		90 18.3682	131 17.8524	232 17.3109	199 17.0759	
		2.4862294		0.0013	0.0009	0.0010	0.0010	
40	UltraVISTA J095850.10+024719.9	149.7087571		75 18.4555	80 17.9656	123 17.4313	101 17.2116	
		2.7888722		0.0011	0.0008	0.0012	0.0018	

Table 7 – *continued* Standard Stars with VISTA-System ZYJHKs

VFS	IAU name	RA ^o Decl. ^o	n Z	Z n Y ± mag	Y n J ± mag	J n H ± mag	H n Ks ± mag	Ks ± mag
41	UltraVISTA J095853.62+024504.2	149.7234241 2.7511805		75 19.2576 0.0012	80 18.6410 0.0012	123 18.1390 0.0019	101 17.8376 0.0021	
42	UltraVISTA J095854.36+023524.9	149.7264829 2.5902461		140 18.3791 0.0013	150 17.8180 0.0007	236 17.2913 0.0011	203 17.0327 0.0012	
43	UltraVISTA J095855.38+024424.0	149.7307359 2.7399894		75 18.7503 0.0010	80 18.1943 0.0008	123 17.6766 0.0013	101 17.4237 0.0019	
44	UltraVISTA J095855.53+024134.4	149.7313560 2.6928870		143 18.0241 0.0008	158 17.5590 0.0006	252 17.0082 0.0009	206 16.7933 0.0009	
45	UltraVISTA J095855.67+024357.8	149.7319650 2.7327131		75 19.8434 0.0022	80 19.2441 0.0020	123 18.7651 0.0025	101 18.4652 0.0033	
46	UltraVISTA J095855.97+024741.0	149.7332239 2.7947267		71 18.4389 0.0009	78 17.8730 0.0008	120 17.4382 0.0013	97 17.1655 0.0016	
47	UltraVISTA J095856.54+021913.8	149.7355822 2.3205009		140 18.8246 0.0010	150 18.3514 0.0007	236 17.7820 0.0012	203 17.5625 0.0014	
48	UltraVISTA J095856.65+022914.1	149.7360512 2.4872598		147 18.0802 0.0008	152 17.6720 0.0006	232 17.0809 0.0009	199 16.9594 0.0010	
49	UltraVISTA J095859.43+024338.2	149.7476065 2.7272746		75 19.1804 0.0016	80 18.6534 0.0014	123 18.1630 0.0017	101 17.9131 0.0027	
50	UltraVISTA J095901.56+024340.8	149.7564883 2.7279976		75 17.9949 0.0011	80 17.4583 0.0008	123 16.8964 0.0010	101 16.6443 0.0012	
51	UltraVISTA J095902.34+024657.9	149.7597522 2.7827541		75 18.4764 0.0012	80 17.8232 0.0009	123 17.2872 0.0014	101 16.9664 0.0015	
52	UltraVISTA J095905.83+024007.4	149.7742789 2.6687115		79 18.2324 0.0011	104 17.8107 0.0008	251 17.2140 0.0010	206 17.0529 0.0010	
53	UltraVISTA J095906.54+024642.1	149.7772394 2.7783479		75 18.8640 0.0011	80 18.3290 0.0011	123 17.7934 0.0014	101 17.5463 0.0018	
54	UltraVISTA J095910.16+023459.8	149.7923194 2.5832843		140 18.1225 0.0012	150 17.5830 0.0006	237 17.1304 0.0011	203 16.8712 0.0010	
55	UltraVISTA J095915.75+024611.1	149.8156044 2.7697500		75 18.2690 0.0009	80 17.7772 0.0010	123 17.2745 0.0010	101 17.0546 0.0013	
56	UltraVISTA J095916.13+024618.3	149.8172140 2.7717426		75 18.2823 0.0009	80 17.8016 0.0008	123 17.2609 0.0013	101 17.0372 0.0014	
57	UltraVISTA J095918.04+024539.0	149.8251519 2.7608222		75 18.3345 0.0010	80 17.8417 0.0007	123 17.3732 0.0012	101 17.1457 0.0014	
58	UCAL J182653.41+040331.5	276.7225557 4.0587376	133 17.4972	131 17.2159 0.0081	156 16.6992 0.0042	157 16.1551 0.0036	167 15.9994 0.0034	
59	UCAL J182718.74+035838.7	276.8280740 3.9774041	140 17.2663	138 17.0325 0.0074	158 16.5879 0.0035	157 16.1728 0.0036	169 16.0038 0.0033	
60	UCAL J182857.59+040542.8	277.2399424 4.0952285	135 17.5724	137 17.2598 0.0085	156 16.7456 0.0041	154 16.3002 0.0037	168 16.0376 0.0037	
61	UCAL J182904.80+043119.7	277.2699877 4.5221425	135 17.6120	133 17.3401 0.0081	157 16.7872 0.0040	156 16.3344 0.0037	167 16.1180 0.0056	
62	UDXS J221217.08+004949.3	333.0711763 0.8303505			26 17.3587 0.0027		22 16.5811 0.0046	
63	UDXS J221217.17+005458.4	333.0715557 0.9162355			26 18.1551 0.0035		22 17.3931 0.0054	
64	UDXS J221217.24+011120.2	333.0718253 1.1889522			26 18.2026 0.0042		22 17.4465 0.0059	
65	UDXS J221308.15-002438.1	333.2839533 -0.4105822			24 17.7879 0.0030		27 17.0097 0.0048	
66	UDXS J221310.28+004403.5	333.2928434 0.7343005			24 17.4070 0.0028	29 16.9648 0.0033	24 16.6916 0.0048	
67	UDXS J221312.69+004516.1	333.3028790 0.7544734			24 17.5557 0.0035	29 17.1132 0.0037	24 16.8090 0.0054	
68	UDXS J221326.27+004504.4	333.3594622 0.7512320			24 18.3002 0.0034	29 17.8259 0.0034	24 17.5602 0.0057	
69	UDXS J221331.78+004432.8	333.3824122 0.7424491			24 17.8077 0.0028	28 17.3309 0.0037	24 16.9979 0.0048	
70	UDXS J221348.36+004340.2	333.4515020 0.7278254			24 17.6807 0.0044	29 17.2226 0.0032	24 16.9545 0.0049	
71	UDXS J221359.70+004405.8	333.4987510 0.7349581			24 17.9313 0.0034	29 17.4356 0.0043	24 17.1440 0.0052	
72	UDXS J221450.27+003128.5	333.7094552 0.5245924			25 17.2650 0.0024	27 16.6699 0.0032	22 16.5297 0.0064	
73	UDXS J221544.78-005649.9	333.9366041 -0.9471891			24 17.3086 0.0024		24 16.4942 0.0050	
74	UDXS J221835.85+012255.2	334.6493784 1.3819891			33 17.8448 0.0039	20 17.2731 0.0054	44 16.9818 0.0051	
75	UDXS J221916.37+002334.2	334.8182190 0.3928450			32 17.6622 0.0034		34 16.8747 0.0054	
76	UDXS J221942.03+004338.8	334.9251383 0.7274321			31 17.4172 0.0030	21 16.8889 0.0016	34 16.7754 0.0047	
77	UDXS J222007.27+002643.6	335.0303119 0.4454373			29 17.4399 0.0040	22 17.0128 0.0035	30 16.9227 0.0049	
78	UDXS J222153.93+003033.1	335.4746997 0.5091841			45 17.3604 0.0031	33 16.7993 0.0034	46 16.5495 0.0049	
79	UDXS J222205.63+001714.8	335.5234567 0.2874459			31 17.1721 0.0036	26 16.6742 0.0036	32 16.5423 0.0052	
80	UDXS J222219.75-000839.9	335.5823067 -0.1444080			28 17.3728 0.0031		33 16.6027 0.0042	
81	UDXS J222227.61+003032.8	335.6150576 0.5091194			31 17.2914 0.0027	26 16.8690 0.0036	32 16.7655 0.0054	

APPENDIX A: SQL QUERIES

SQL for the DXS

```

SELECT s.ra, s.dec, s.jAperMag3, v.jMeanMag, v.jnGoodObs,
s.jAperMag3err, v.jMagRms, v.jMagMAD,
s.hAperMag3, v.hMeanMag, v.hnGoodObs,
s.hAperMag3err, v.hMagRms, v.hMagMAD, s.kAperMag3,
v.kMeanMag, v.knGoodObs, s.kAperMag3err, v.kMagRms,
v.kMagMAD FROM dxsSource AS s, dxsVariability AS v WHERE s.sourceID=v.sourceID AND
s.mergedClass=-1 AND v.variableClass=0 AND s.dec
> -30 AND v.kMeanMag>16.5 AND v.jnGoodObs>20
AND v.hnGoodObs>5 AND v.knGoodObs>20 AND
(v.jMagMAD/(SQRT(v.jnGoodObs - 1)))<=0.004
AND (v.hMagMAD/(SQRT(v.hnGoodObs - 1)))<=0.004
AND (v.kMagMAD/(SQRT(v.knGoodObs - 1)))<=0.004
AND (v.jMagRms/SQRT(v.jnGoodObs - 1))<=0.006
AND (v.hMagRms/SQRT(v.hnGoodObs - 1))<=0.006
AND (v.kMagRms/SQRT(v.knGoodObs - 1))<=0.006
AND ((s.jAperMag3 - v.jMeanMag) < 2.5 *
(SQRT(s.jAperMag3err * s.jAperMag3err
+ (v.jMagRms/SQRT(v.jnGoodObs - 1)) *
(v.jMagRms/SQRT(v.jnGoodObs - 1))))))
AND ((s.hAperMag3 - v.hMeanMag) < 2.5 *
(SQRT(s.hAperMag3err * s.hAperMag3err
+ (v.hMagRms/SQRT(v.hnGoodObs - 1)) *
(v.hMagRms/SQRT(v.hnGoodObs - 1))))))
AND ((s.kAperMag3 - v.kMeanMag) < 2.5 *
(SQRT(s.kAperMag3err * s.kAperMag3err
+ (v.kMagRms/SQRT(v.knGoodObs - 1)) *
(v.kMagRms/SQRT(v.knGoodObs - 1))))))

```

SQL for the UDS

NOTE: there is an offset between AperMag3 and MeanMag due to the former not being aperture-corrected.

```

SELECT s.ra, s.dec, s.jAperMag3, v.jMeanMag, v.jnGoodObs,
s.jAperMag3err, v.jMagRms, v.jMagMAD,
s.hAperMag3, v.hMeanMag, v.hnGoodObs,
s.hAperMag3err, v.hMagRms, v.hMagMAD, s.kAperMag3,
v.kMeanMag, v.knGoodObs, s.kAperMag3err, v.kMagRms,
v.kMagMAD FROM udsSource AS s, udsVariability AS v WHERE s.sourceID=v.sourceID AND s.mergedClass=-1 AND v.variableClass=0 AND v.kMeanMag>17.5
AND v.jnGoodObs>100 AND v.hnGoodObs>100 AND
v.knGoodObs>100 AND (v.jMagMAD/(SQRT(v.jnGoodObs
- 1)))<=0.004 AND (v.hMagMAD/(SQRT(v.hnGoodObs
- 1)))<=0.004 AND (v.kMagMAD/(SQRT(v.knGoodObs
- 1)))<=0.004 AND (v.jMagRms/SQRT(v.jnGoodObs
- 1))<=0.006 AND (v.hMagRms/SQRT(v.hnGoodObs
- 1))<=0.006 AND (v.kMagRms/SQRT(v.knGoodObs
- 1))<=0.006 AND s.jAperMag3err<=0.003 AND
s.hAperMag3err<=0.003 AND s.kAperMag3err<=0.003
AND (s.jAperMag3 - v.jMeanMag < 0.20) AND
(s.jAperMag3 - v.jMeanMag > 0.18) AND (s.hAperMag3
- v.hMeanMag < 0.21) AND (s.hAperMag3 - v.hMeanMag
> 0.19) AND (s.kAperMag3 - v.kMeanMag < 0.19) AND
(s.kAperMag3 - v.kMeanMag > 0.17)

```

SQL for the UltraVISTA

NOTE: there is an offset between AperMag3 and MeanMag

due to differences in the aperture correction and zeropoints.

```

SELECT v.sourceID, s.ra, s.dec, v.variableClass,
s.yclassStat, s.jclassStat, s.hclassStat,
s.ksclassStat, s.yAperMag3, v.yMeanMag,
v.ynGoodObs, s.yAperMag3err, v.yMagRms,
v.yMagMAD, s.jAperMag3, v.jMeanMag, v.jnGoodObs,
s.jAperMag3err, v.jMagRms, v.jMagMAD, s.hAperMag3,
v.hMeanMag, v.hnGoodObs, s.hAperMag3err,
v.hMagRms, v.hMagMAD, s.ksAperMag3, v.ksMeanMag,
v.ksnGoodObs, s.ksAperMag3err, v.ksMagRms,
v.ksMagMAD FROM ultravistaSource AS s, ultravistaVariability AS v WHERE s.sourceID=v.sourceID AND v.variableClass=0 AND v.ksMeanMag > 16.5
AND s.hclassStat>0.7 AND s.ksclassStat>0.7
AND v.ynGoodObs>=20 AND v.jnGoodObs>=20 AND
v.hnGoodObs>=20 AND v.ksnGoodObs>=20 AND
(v.yMagMAD/(SQRT(v.ynGoodObs- 1)))<=0.004 AND
(v.jMagMAD/(SQRT(v.jnGoodObs- 1)))<=0.004 AND
(v.hMagMAD/(SQRT(v.hnGoodObs - 1)))<=0.004 AND
(v.ksMagMAD/(SQRT(v.ksnGoodObs- 1)))<=0.004
AND (v.yMagRms/SQRT(v.ynGoodObs- 1))<=0.006
AND (v.jMagRms/SQRT(v.jnGoodObs- 1))<=0.006
AND (v.hMagRms/SQRT(v.hnGoodObs- 1))<=0.006
AND (v.ksMagRms/SQRT(v.ksnGoodObs- 1))<=0.006
AND ((s.yAperMag3 - v.yMeanMag) <= 0.0) AND
((s.yAperMag3 - v.yMeanMag) >= -0.04) AND
((s.jAperMag3 - v.jMeanMag) <= -0.04) AND
((s.jAperMag3 - v.jMeanMag) >= -0.08) AND
((s.hAperMag3 - v.hMeanMag) <= 0.04) AND
((s.hAperMag3 - v.hMeanMag) >= 0.00) AND
((s.ksAperMag3 - v.ksMeanMag) <= 0.02) AND
((s.ksAperMag3 - v.ksMeanMag) >= -0.02)

```

SQL for VIDEO

```

SELECT v.sourceID, s.ra, s.dec, s.mergedClass,
s.zAperMag3, v.zMeanMag, v.znGoodObs,
s.zAperMag3err, v.zMagRms, v.zMagMAD,
s.yAperMag3, v.yMeanMag, v.ynGoodObs, s.yAperMag3err,
v.yMagRms, v.yMagMAD, s.jAperMag3, v.jMeanMag, v.jnGoodObs,
s.jAperMag3err, v.jMagRms, v.jMagMAD,
s.hAperMag3, v.hMeanMag, v.hnGoodObs,
s.hAperMag3err, v.hMagRms, v.hMagMAD,
s.ksAperMag3, v.ksMeanMag,
v.ksnGoodObs, s.ksAperMag3err, v.ksMagRms,
v.ksMagMAD FROM videoSource AS s, videoVariability AS v WHERE s.sourceID=v.sourceID AND s.mergedClass in (-1, -2) AND v.variableClass=0 AND s.dec > -30 AND v.ksMeanMag > 17.5 AND v.ynGoodObs>20
AND v.jnGoodObs> 20 AND v.hnGoodObs > 20 AND
v.ksnGoodObs > 20 AND (v.yMagMAD/(SQRT(v.ynGoodObs
- 1)))<=0.004 AND (v.jMagMAD/(SQRT(v.jnGoodObs
- 1)))<=0.004 AND (v.hMagMAD/(SQRT(v.hnGoodObs -
1)))<=0.004 AND (v.ksMagMAD/(SQRT(v.ksnGoodObs
- 1)))<=0.004 AND (v.yMagRms/SQRT(v.ynGoodObs
- 1))<=0.006 AND (v.jMagRms/SQRT(v.jnGoodObs
- 1))<=0.006 AND (v.hMagRms/SQRT(v.hnGoodObs
- 1))<=0.006 AND (v.ksMagRms/SQRT(v.ksnGoodObs
- 1))<=0.006 AND ((s.yAperMag3 - v.yMeanMag)
< 2.5 * (SQRT(s.yAperMag3err * s.yAperMag3err
+ (v.yMagRms/SQRT(v.ynGoodObs - 1)) *
(v.yMagRms/SQRT(v.ynGoodObs - 1))))))

```

```

AND ((s.jAperMag3 - v.jMeanMag) < 2.5 *
(SQRT(s.jAperMag3err * s.jAperMag3err
+ (v.jMagRms/SQRT(jnGoodObs - 1)) *
(v.jMagRms/SQRT(v.jnGoodObs - 1))))))
AND ((s.hAperMag3 - v.hMeanMag) < 2.5 *
(SQRT(s.hAperMag3err * s.hAperMag3err
+ (v.hMagRms/SQRT(v.hnGoodObs - 1)) *
(v.hMagRms/SQRT(v.hnGoodObs - 1)))))) AND
((s.ksAperMag3 - v.ksMeanMag) < 2.5 *
(SQRT(s.ksAperMag3err * s.ksAperMag3err
+ (v.ksMagRms/SQRT(ksnGoodObs - 1)) *
(v.ksMagRms/SQRT(ksnGoodObs - 1))))))

```

SQL for WFCAMCAL

```

SELECT s.ra, s.dec, s.zAperMag3, v.zMeanMag, v.znGoodObs,
s.zAperMag3err, v.zMagRms, v.zMagMAD,
s.yAperMag3, v.yMeanMag, v.ynGoodObs,
s.yAperMag3err, v.yMagRms, v.yMagMAD, s.jAperMag3,
v.jMeanMag, v.jnGoodObs, s.jAperMag3err, v.jMagRms,
v.jMagMAD, s.hAperMag3, v.hMeanMag, v.hnGoodObs,
s.hAperMag3err, v.hMagRms, v.hMagMAD, s.kAperMag3,
v.kMeanMag, v.knGoodObs, s.kAperMag3err,
v.kMagRms, v.kMagMAD FROM calSource AS s, cal-
Variability AS v WHERE s.sourceID=v.sourceID
AND s.mergedClass=-1 AND v.variableClass=0
AND v.kMeanMag>16 AND v.ynGoodObs>100 AND
v.jnGoodObs>100 AND v.hnGoodObs>100 AND
v.knGoodObs>100 AND (v.yMagMAD/(SQRT(v.ynGoodObs
- 1)))<=0.004 AND (v.jMagMAD/(SQRT(v.jnGoodObs
- 1)))<=0.004 AND (v.hMagMAD/(SQRT(v.hnGoodObs
- 1)))<=0.004 AND (v.kMagMAD/(SQRT(v.knGoodObs
- 1)))<=0.004 AND (v.yMagRms/SQRT(v.ynGoodObs
- 1))<=0.006 AND (v.jMagRms/SQRT(v.jnGoodObs
- 1))<=0.006 AND (v.hMagRms/SQRT(v.hnGoodObs
- 1))<=0.006 AND (v.kMagRms/SQRT(v.knGoodObs
- 1))<=0.006 AND ((s.yAperMag3 - v.yMeanMag)
< 2.5 * (SQRT(s.yAperMag3err * s.yAperMag3err
+ (v.yMagRms/SQRT(v.ynGoodObs - 1)) *
(v.yMagRms/SQRT(v.ynGoodObs - 1))))))
AND ((s.jAperMag3 - v.jMeanMag) < 2.5 *
(SQRT(s.jAperMag3err * s.jAperMag3err
+ (v.jMagRms/SQRT(v.jnGoodObs - 1)) *
(v.jMagRms/SQRT(v.jnGoodObs - 1))))))
AND ((s.hAperMag3 - v.hMeanMag) < 2.5 *
(SQRT(s.hAperMag3err * s.hAperMag3err
+ (v.hMagRms/SQRT(v.hnGoodObs - 1)) *
(v.hMagRms/SQRT(v.hnGoodObs - 1))))))
AND ((s.kAperMag3 - v.kMeanMag) < 2.5 *
(SQRT(s.kAperMag3err * s.kAperMag3err
+ (v.kMagRms/SQRT(v.knGoodObs - 1)) *
(v.kMagRms/SQRT(v.knGoodObs - 1))))))

```

APPENDIX B: POSSIBLE GALAXIES

Table B1. Data for Sources Identified as Possible Galaxies

IR Survey Name	Flag	RA ^o Decl. ^o	RA ^o ± mas	Decl. ^o ± mas	Gaia Parallax ± mas	μ RA ± mas yr ⁻¹	μ Decl. ± mas yr ⁻¹	G	Pan-STARRS g ±	meanPSF r ±	AB mags i ±	z ±	y Z ± ±	VISTA Y ±	Vega mags J ±	H ±	Ks ±	
UDS 34.0387095	1	34.038750344750	-5.393224535432	2.0828	4.0926	-8.9745	20.7447	21.9886	20.8532	20.0351	19.6112	19.3296		18.2804	17.7371	17.5162		
-5.3931529		1.43711	1.25846	1.7445	2.8440	2.8502		0.0532	0.0207	0.0182	0.0205	0.0493		0.0020	0.0018	0.0032		
UDS 34.0632333	2													18.5751	18.0561	17.7469		
-5.4123534										0.0323	0.0234	0.0550		0.0022	0.0022	0.0034		
UDS 34.0725841	1	34.072628552341	-5.367278457759	0.8299	-0.2061	-4.3827	19.6567	19.9357	19.6258	19.5138	19.4769	19.4584		18.6107	18.3250	18.2648		
-5.3672215		0.62595	0.50292	0.8537	1.0136	0.8118		0.0124	0.0226	0.0115	0.0154	0.0358		0.0017	0.0021	0.0027		
UDS 34.2131217	1	34.213113369956	-4.741901511918	0.2797	-0.3908	-3.9373	19.0469	19.3203	19.0049	18.8999	18.9062	18.9296		18.0473	17.7772	17.6997		
-4.7418757		0.37411	0.26570	0.4289	0.5696	0.4495		0.0056	0.0104	0.0060	0.0124	0.0333		0.0015	0.0018	0.0026		
UDS 34.2836251	1							21.3821	20.3480	19.9655	19.5100			18.4806	17.9502	17.6859		
-5.5122054								0.0442	0.0143	0.0380	0.1018			0.0021	0.0019	0.0034		
UDS 34.3328267	1							21.9449	20.7901	20.1041	19.6980			18.5164	17.9305	17.6470		
-5.2005163								0.1826	0.0256	0.0299	0.0670			0.0024	0.0021	0.0036		
UDS 34.3905494	1							21.5139	20.3717	19.8806	19.4405			18.3636	17.8970	17.6138		
-5.1081487								0.0772	0.0156	0.0365	0.0262			0.0022	0.0023	0.0033		
UDS 34.4226317	2	34.422696476887	-5.319394990941				20.9700	21.9304	21.1561	20.1452	19.6900	19.4092		18.3216	17.8349	17.5908		
-5.3193711		2.88829	7.68729					0.1454	0.0364	0.0123	0.0222	0.0706		0.0022	0.0019	0.0033		
UDS 34.4476113	1	34.447624608681	-5.115979177256	-0.4065	-0.5852	-1.9265	19.8064	20.2081	19.7500	19.5395	19.4742	19.3670		18.4760	18.0645	17.9778		
-5.1159795		0.65716	0.53976	1.0029	0.8452	0.8947		0.0176	0.0209	0.0161	0.0194	0.0826		0.0021	0.0024	0.0032		
UDS 34.4614311	1							21.7428	20.6132	20.0122	19.5653			18.4719	17.9032	17.6350		
-5.1427924								0.0493	0.0315	0.0264	0.0470			0.0023	0.0022	0.0035		
UDS 34.5115289	2							21.8823	20.7066	20.0423	19.6703			18.4488	18.0040	17.6905		
-5.2450022								0.1836	0.0242	0.0221	0.0723			0.0023	0.0022	0.0033		
UDS 34.5259644	1,2							21.8031	21.7587	20.5598	19.8926	19.4451		18.3814	17.9104	17.6043		
-5.2628373								0.1946	0.0513	0.0334	0.0416	0.0253		0.0022	0.0021	0.0033		
UDS 34.5434314	1	34.543463866893	-5.171139515300	0.6485	-0.1422	-2.7368	19.5637	20.1365	19.4735	19.2855	19.1939	19.1399		18.1998	17.8036	17.7009		
-5.1711112		0.50344	0.35905	0.5815	0.7438	0.6398		0.0043	0.0370	0.0119	0.0140	0.0194		0.0016	0.0017	0.0028		
UDS 34.5471946	1	34.547230921295	-5.176806400064	0.3174	6.3978	-9.4345	19.6890	20.1476	19.6217	19.3895	19.2849	19.1582		18.2589	17.8400	17.7102		
-5.1767668		0.55339	0.39306	0.6321	0.8295	0.7060		0.0146	0.0144	0.0098	0.0106	0.0339		0.0018	0.0017	0.0029		
UDS 34.5487686	1							21.3331	20.3608	19.9194	19.4469			18.4316	17.9402	17.7101		
-5.3929635								0.0609	0.0228	0.0202	0.0669			0.0021	0.0019	0.0032		
UDS 34.5590572	1	34.559105809238	-4.927423298600	-0.1536	3.8946	-3.5982	19.2729	19.5643	19.2749	19.1408	19.1434	19.1397		18.2892	18.0363	17.9828		
-4.9273779		0.60308	0.45118	0.7568	0.9018	0.7822		0.0208	0.0020	0.0108	0.0169	0.0479		0.0014	0.0020	0.0027		
UDS 34.6711628	1							22.0633	20.8571	20.2795	19.8136			18.7770	18.3086	18.0321		
-5.1675618								0.1102	0.0530	0.0278	0.0633			0.0023	0.0023	0.0035		
UDS 34.6858172	1	34.685894904232	-4.703061471509	0.3458	11.8361	-6.0594	19.4035	19.7527	19.3522	19.1732	19.1381	19.0562		18.1939	17.8580	17.7599		
-4.7030450		0.63685	0.48470	0.9393	0.8860	0.7466		0.0115	0.0128	0.0096	0.0110	0.0295		0.0016	0.0017	0.0027		
UDS 34.7683910	2							21.8062	20.8612	20.2771	19.8818			18.7346	18.2226	17.9366		
-5.0398575								0.0160	0.0548	0.0322	0.0947			0.0022	0.0021	0.0034		
UDS 34.7885620	1,2							22.1346	21.3754	20.3912	19.8497	19.4869		18.4368	17.9835	17.7201		
-4.8819019								0.1739	0.0588	0.0258	0.0361	0.0416		0.0020	0.0018	0.0033		
UDS 34.8300954	1							21.9653	21.7962	20.5947	20.0403	19.6512		18.5714	18.1156	17.8530		
-5.0119407								0.1578	0.0556	0.0144	0.0215	0.0364		0.0021	0.0020	0.0032		
UDS 34.8397670	2							22.6037	21.9676	20.7494	20.0257	19.6058		18.3350	17.8180	17.5136		
-4.7351863								0.2478	0.0095	0.0279	0.0278	0.0594		0.0022	0.0018	0.0034		
UDS 34.8474938	1	34.847507314304	-5.335409824465				20.6009	21.8297	20.5566	19.9651	19.6698	19.3245		18.3256	17.7770	17.5709		
-5.3353341		2.19216	2.17627					0.0467	0.0499	0.0117	0.0189	0.0746		0.0021	0.0019	0.0033		
UDS 34.8707629	1							21.2460	20.3159	19.6511				18.4190	17.9452	17.5810		
-5.1977157								0.0373	0.0206	0.0313				0.0022	0.0020	0.0035		
UDS 34.8890054	2	34.889028664017	-4.822057442111				21.0061	22.1342	21.2954	20.2237	19.7842	19.5169		18.3800	17.8426	17.6186		
-4.8219876		3.44884	10.04841					0.1309	0.0398	0.0166	0.0195	0.0354		0.0023	0.0022	0.0035		
VIDEO 35.0364084	2							21.0837	20.2244	19.7591			18.9637	18.3366	17.8180	17.5239		
-4.9878868								0.0253	0.0486	0.0730			0.0022	0.0017	0.0028	0.0039		
VIDEO 35.0634277	1							20.8514	20.2694	19.7768			19.3151	18.7851	18.1603	17.9089		
-4.8899093								0.0269	0.0319	0.0136			0.0023	0.0025	0.0032	0.0040		
VIDEO 35.1404271	1,2							21.9319	21.0028	20.1999	19.8095	19.3154		18.8489	18.3912	17.9152	17.7179	
-4.5489657								0.0104	0.0309	0.0243	0.0215	0.0698		0.0028	0.0024	0.0025	0.0050	
VIDEO 35.1942372	1	35.194280993865	-4.866190049713		15.1823	-5.4891	20.0758	20.6159	19.9907	19.7274	19.6449	19.4795		18.9034	18.5732	18.1256	18.0160	
-4.8661564		0.68076	0.48087		1.0299	0.8569		0.0223	0.0123	0.0141	0.0150	0.0444		0.0015	0.0030	0.0035	0.0045	
VIDEO 35.2729996	1							22.3620	21.9127	20.8624	20.4448	20.0550		19.4404	18.9522	18.5055	18.2381	
-5.0046632								0.2410	0.1926	0.0450	0.0543	0.0634		0.0021	0.0030	0.0054	0.0047	
VIDEO 35.3783901	1							21.7415	20.8513	20.2565	19.7882			19.2379	18.7073	18.2273	17.9349	
-5.2565457								0.0291	0.0190	0.0312	0.1079			0.0022	0.0026	0.0051	0.0045	
VIDEO 35.3829208	1							21.8337	20.9823	20.4489	20.1954	19.8148		19.3882	18.9576	18.3868	18.1676	
-4.9597282								0.0144	0.0513	0.0243	0.0340	0.1005		0.0023	0.0030	0.0048	0.0056	
VIDEO 35.3846956	1							21.7485	20.7684	20.2319	19.8866			19.2649	18.7794	18.3458	18.0929	
-4.4494825								0.1597	0.0302	0.0201	0.0325			0.0019	0.0022	0.0044	0.0024	
VIDEO 35.4517777	1	35.451819126350	-4.484091070150		1.9364	-9.6414	20.6038	21.3819	20.5229	20.1307	19.8683	19.7105		19.1223	18.7165	18.1694	18.0295	
-4.4840691		0.94865	1.11158		2.0494	2.9685		0.0351	0.0226	0.0149	0.0245	0.0322		0.0025	0.0027	0.0029	0.0044	
VIDEO 35.5506321	1							21.7361	20.9951	20.5266	20.0104			19.4807	18.9698	18.4913	18.2201	
-4.6345957								0.1500	0.0219	0.0151	0.0008			0.0022	0.0024	0.0045	0.0047	
VIDEO 35.5527346	1							20.9396	20.2919	19.8182				19.2588	18.7094	18.2230	17.9311	
-4.9423113								0.0302	0.0286	0.0443				0.0031	0.0024	0.0049	0.0053	
VIDEO 35.6795871	1,2							21.4877	20.8680	20.5088	20.1017			19.6523	19.2109	18.7295	18.4770	
-5.2156567								0.072										

Table B1 – *continued* Data for Sources Identified as Possible Galaxies

IR Survey Name	RA ^o Decl. ^o	Flag	RA ^o ± mas	Gaia				G	Pan-STARRS meanPSF AB mags					VISTA Vega mags				
				Decl. ^o ± mas	Parallax ± mas	μ RA ± mas yr ⁻¹	μ Decl. ± mas yr ⁻¹		<i>g</i> ±	<i>r</i> ±	<i>i</i> ±	<i>z</i> ±	<i>y</i> ±	<i>Z</i> ±	<i>Y</i> ±	<i>J</i> ±	<i>H</i> ±	<i>Ks</i> ±
VIDEO	35.7408174 -5.2873244	1							21.5238	20.6134	20.1719	19.8212		19.2800	18.8185	18.3103	18.0684	
VIDEO	35.7414119 -4.8313848	1							0.0496	0.0220	0.0486	0.0911		0.0020	0.0034	0.0036	0.0044	
VIDEO	35.7655027 -5.0096975	2							21.6836	20.7530	20.1849	19.7888		19.2312	18.7566	18.2657	18.0543	
VIDEO	35.8265344 -4.9284675	1							0.0472	0.0258	0.0118	0.0540		0.0028	0.0027	0.0054	0.0052	0.0044
VIDEO	35.9375707 -4.4698025	1							21.2297	21.2518	20.2632	19.7629	19.3965	18.7797	18.3147	17.8308	17.6399	
VIDEO	35.9460049 -4.5784901	1							0.0982	0.0172	0.0159	0.0173	0.0239	0.0023	0.0027	0.0022	0.0044	
VIDEO	36.0710161 -4.4139791	1	36.071091123688 1.28068	-4.413983619399 0.78531		9.7112 2.0081	5.4933 1.8851	20.5758	21.6183	20.5739	19.9287	19.5959	19.3285	19.0171	18.7109	18.2809	17.7485	17.5633
VIDEO	36.3757309 -4.0906468	1	36.375784685588 0.55848	-4.090708837809 0.6330	0.7755 0.6330	2.5459 1.0837	-12.1766 0.7780	19.6292	20.0069	19.5620	19.3690	19.3078	19.2324	18.7250	18.5751	18.2832	17.9134	17.8111
VIDEO	36.6440048 -4.8126421	1,2	36.644038728037 1.46427	-4.812678556252 2.36397				20.7420	21.7597	20.8992	20.1040	19.6922	19.4314	19.1277	18.8228	18.3785	17.8625	17.6593
VIDEO	36.6880308 -4.8435704	1							0.0488	0.0123	0.0094	0.0117	0.0428	0.0020	0.0013	0.0019	0.0033	0.0035
VIDEO	36.7097084 -4.5243931	1	36.709737717167 0.67107	-4.524403538704 0.72915	-1.0689 0.7793	1.9018 1.5715	-6.7427 2.0535	20.1873	20.5807	20.0892	19.8732	19.7632	19.6094	19.2129	19.0658	18.7609	18.3770	18.2744
VIDEO	52.0097291 -27.6189720	1							0.0185	0.0116	0.0072	0.0237	0.0367	0.0016	0.0019	0.0026	0.0043	0.0049
VIDEO	52.4462564 -27.8237538	1								20.6892	20.0223	19.5554	19.3653	18.9275	18.4124	18.0183	17.7420	17.6049
VIDEO	52.7557893 -27.4488061	1	52.755781300291 0.27729	-27.448804052908 0.43118	-0.0858 0.5462	0.9688 0.6486	-6.4672 0.8882	20.1116	20.8253	20.0564	19.6631	19.5744	19.3967	18.9411	18.7801	18.4311	17.9267	17.8077
VIDEO	52.9274555 -27.4700603	2							0.0423	0.0191	0.0086	0.0158	0.0304	0.0040	0.0024	0.0033	0.0032	0.0056
VIDEO	53.2112772 -27.2074247	1	53.211345751370 0.81033	-27.207439654263 1.24769	0.7372 1.6953	18.1557 1.7435	-9.0152 2.4402	20.8317	20.9198	20.1251	19.7043	19.5700	19.0776	18.8092	18.3537	17.9459	17.6940	
VIDEO	53.3409644 -27.8449838	1,2	53.340983559831 0.51344	-27.845025165778 0.97876	-0.7650 1.1896	12.5581 1.3332	-17.2877 2.4270	20.8951	21.8998	20.8341	20.2670	19.9444	19.6204	19.3216	19.0852	18.6317	18.1686	17.9531
CAL	87.6810876 16.3722372	1	87.681107197032 0.22657	16.372256410953 0.19663	0.5648 0.2468	0.7724 0.4921	-2.9455 0.4134	18.3802	19.0977	18.3361	17.9998	17.8359	17.6862	17.2938	17.1385	16.6670	16.2310	16.0786
CAL	87.7003018 15.8245989	1	87.700299075751 0.14842	15.824548535954 0.13937	0.2557 0.1558	2.3478 0.3193	-1.9968 0.2695	17.9648	18.5295	17.9302	17.6653	17.5282	17.4332	17.0166	16.8713	16.4748	16.1133	15.9915
UVISTA	149.3964243 1.6279765	1	149.396426037047 0.57794	1.627980712026 0.98644				18.0515	18.3400	18.0239	17.9150	17.8861	17.8455	17.2622	17.0280	16.7440	16.6915	
UVISTA	149.7445300 2.7571103	1							0.0118	0.0064	0.0055	0.0105	0.0158	0.0018	0.0016	0.0024	0.0041	
UVISTA	149.7772837 2.7455382	1								22.7023	21.1733	20.5041	19.9158	19.3617	18.8135	18.2684	18.0102	
UVISTA	149.7833198 2.7839207	1,2								0.0249	0.0378	0.0375	0.0424	0.0012	0.0015	0.0018	0.0025	
UVISTA	149.7947682 2.7653208	2	149.794759440482 3.37503	2.765319258775 5.25135				20.7674	21.7339	20.6624	20.1608	19.8799	19.7512	19.0779	18.6551	18.1258	17.9319	
UVISTA	149.8276357 2.7542743	2							0.0749	0.0210	0.0212	0.0273	0.0567	0.0012	0.0013	0.0018	0.0027	
CAL	312.6852216 6.1850892	1	312.685178912608 0.17048	6.185061145961 0.12374	-0.5758 0.2360	-10.5551 0.3446	-9.0282	18.1012	18.5962	18.0192	17.7478	17.6417	17.5611	17.1128	16.9912	16.6060	16.1981	16.0897
DXS	333.0678519 0.2463832	1	312.685178912608 0.17048	6.185061145961 0.12374	-0.5758 0.2360	-10.5551 0.3446	-9.0282	18.1012	18.7337	18.3275	18.1735	18.1308	18.0981	0.0069	0.0037	0.0041	0.0035	0.0049
DXS	333.1814666 0.5377071	1	333.071813320305 1.68843	1.188906051043 1.04016	1.2559 1.9038	-0.023 1.6055	-4.0006	20.2384	19.2513	18.8487	18.6976	18.6483	18.5967		0.0035	0.0031	0.0050	
DXS	333.1895155 0.7402520	1	333.181502744943 0.39660	0.537661770120 0.45081	0.0649 0.4259	3.5358 0.6231	-4.0912	18.8806	19.5321	19.0986	18.9363	18.8497	18.8137		0.0048	0.0059	0.0057	
DXS	333.2850037 -0.9465384	1	333.283923388425 1.74142	-0.410674164956 1.61698				20.7836	22.1518	21.4454	20.0104	19.3589	18.8466		17.9090	17.5613	17.4914	
DXS	333.4320262 0.7382165	2							0.1829	0.0420	0.0108	0.0166	0.0378		0.0036	0.0053	0.0055	
DXS	333.6445754 0.5342898	1	333.644571087727 0.25131	0.534256767502 0.26301	0.2749 0.2872	-3.1602 0.4120	-3.8142	18.4809	18.7319	18.4562	18.3525	18.3092	18.2465		18.0916	17.5841	17.2524	
DXS	334.8159877 1.0973756	1	334.815996502933 0.20974	1.097322542853 0.22680	0.0244 0.2655	1.5929 0.5427	-10.8697	18.1228	18.4275	18.1034	17.9505	17.8859	17.8456		0.0051	0.0052	0.0059	
DXS	334.9984584 0.7319926	1,2	334.998484432193 9.42378	0.731938922959 1.34213				21.0597	21.9666	21.3736	20.1511	19.4318	18.9772		17.4536	17.1791	17.1329	
DXS	335.0317979 -0.1094196	2	335.031828446221 2.07486	-0.109437792442 2.32689	-3.5425 2.5997	2.4789 3.6136	7.4609	20.5124	22.1153	21.1476	19.7651	19.0880	18.7638		0.0032	0.0049	0.0055	
DXS	335.1704830 0.5104702	1	335.170502659949 0.29755	0.510427567859 0.28425	0.5604 0.3486	0.1249 0.3769	-8.3115	18.3110	18.7558	18.2612	18.0592	17.9900	17.9099		17.4985	17.0236	16.7454	
DXS	335.3820794 1.1622935	2	335.382075808196 2.11766	1.162262771954 2.10523				20.2360	21.5394	20.2713	19.5524	19.1903	18.9633		0.0032	0.0038	0.0058	
DXS	335.6191436 1.1619136	1	335.619172934901 0.32319	1.161866766328 0.41983	0.1049 0.3448	9.2640 0.5976	-10.4955	18.6493	19.1149	18.6094	18.3917	18.2874	18.2377		17.7985	17.2037	16.9959	
									0.0163	0.0081	0.0038	0.0119	0.0160		0.0029	0.0034	0.0047	

Flag as possible galaxy: (1) *HKYJ* and/or *HKYJ* colour (2) one or more Pan-STARRS filters have Kron magnitudes brighter than the PSF magnitude by 0.05 mag.

**APPENDIX C: STARS EXCLUDED FROM THE
FINAL SAMPLE**

Table C1. Data for Sources Identified as Likely Stars with Atypical Colours

IR Survey Name	RA [°] Decl. [°]	Pan-STARRS meanPSF AB mags					VISTA Vega mags				
		<i>g</i> ± mag	<i>r</i> ± mag	<i>i</i> ± mag	<i>z</i> ± mag	<i>y</i> ± mag	<i>Z</i> ± mag	<i>Y</i> ± mag	<i>J</i> ± mag	<i>H</i> ± mag	<i>K_s</i> ± mag
UDS	34.2123189	21.7954	21.6093	20.3610	19.8250	19.4105			18.3512	17.8466	17.6092
	-4.8947515	0.1793	0.0835	0.0147	0.0315	0.1145			0.0021	0.0019	0.0033
UDS	34.4109253		21.7650	20.7175	19.9836	19.7015			18.5133	18.0228	17.7496
	-5.2700687		0.1675	0.0290	0.0255	0.0731			0.0022	0.0019	0.0033
UDS	34.5890235		21.8035	20.8665	20.1967	19.7849			18.6272	18.1619	17.9030
	-4.9424929		0.0724	0.0345	0.0278	0.0966			0.0021	0.0020	0.0033
VIDEO	35.1128282	21.6591	21.9368	20.8698	20.1845	19.8103		19.1399	18.6294	18.1766	17.9161
	-4.9552726	0.1655	0.1798	0.0538	0.0102	0.0657		0.0020	0.0025	0.0028	0.0024
VIDEO	52.0756099				20.8638	19.9288	20.2966	19.4553	18.6419	18.1248	17.7045
	-27.8683759				0.0682	0.0791	0.0063	0.0043	0.0026	0.0037	0.0050
DXS	333.3096800		21.6666	20.8225	20.1457	19.7224			18.3903	17.8713	17.5302
	0.7372219		0.1489	0.0248	0.0391	0.0516			0.0040	0.0053	0.0051
DXS	333.8609666		21.4346	20.2861	19.5578	19.1589			18.0057	17.4381	17.1374
	0.7496622		0.1123	0.0126	0.0266	0.2435			0.0055	0.0036	0.0047

Table C2. Likely Stars with Deviant Mean and Aperture Magnitudes

Survey	RA [°] Decl. [°]	<i>Z</i> mean ± mag	<i>Z</i> Aper ± mag	<i>Y</i> mean ± mag	<i>Y</i> Aper ± mag	<i>J</i> mean ± mag	<i>J</i> Aper ± mag	<i>H</i> mean ± mag	<i>H</i> Aper ± mag	<i>K</i> (s) mean ± mag	<i>K</i> (s) Aper ± mag
VIDEO	35.1191358			19.2971	19.2615	18.8033	18.8031	18.3092	18.3053	18.0712	18.0743
	-4.4549895			0.0025	0.0093	0.0037	0.0086	0.0035	0.0082	0.0045	0.0090
CAL	87.6786696	17.4104	17.5649	17.2602	17.2153	16.7709	16.7628	16.1723	16.1782	16.0033	16.0350
	15.8554744	0.0039	0.0183	0.0031	0.0153	0.0037	0.0172	0.0034	0.0138	0.0036	0.0211
CAL	276.8783681	17.3634	17.3458	17.0501	17.0870	16.5950	16.6020	16.0855	16.0882	16.0014	15.9022
	4.5214510	0.0031	0.0159	0.0028	0.0144	0.0030	0.0167	0.0027	0.0148	0.0040	0.0218
CAL	277.1729349	17.5923	17.6030	17.3352	17.3172	16.8237	16.8150	16.3529	16.2602	16.1989	16.0637
	4.5807735	0.0035	0.0184	0.0028	0.0159	0.0035	0.0180	0.0031	0.0161	0.0050	0.0242

APPENDIX D: FINDING CHARTS

This paper has been typeset from a $\text{\TeX}/\text{\LaTeX}$ file prepared by the author.

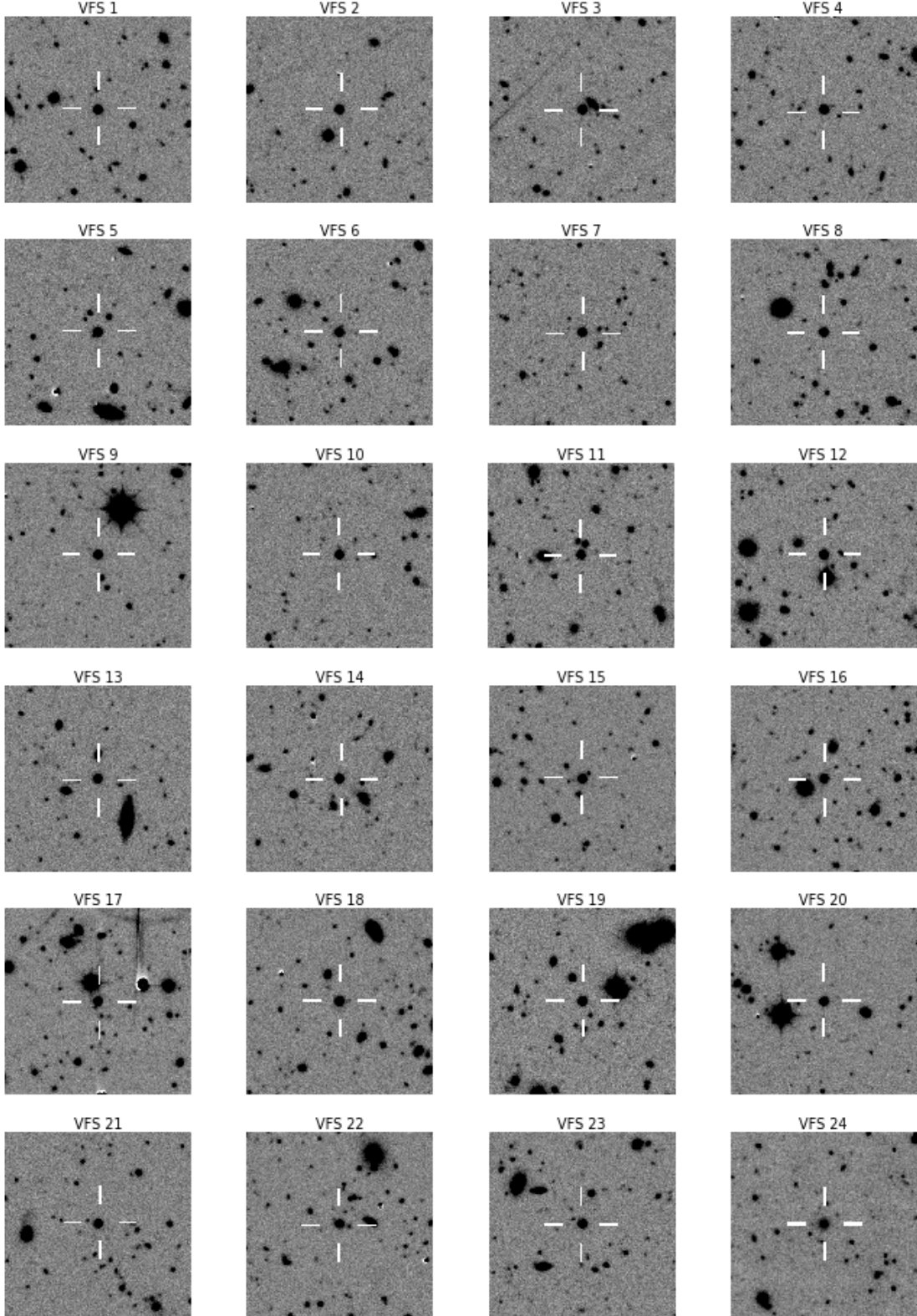


Figure D1. J-band finding charts, 1 arcminute field-of-view, for the standards presented above in Table 7. North is up and east to the left.

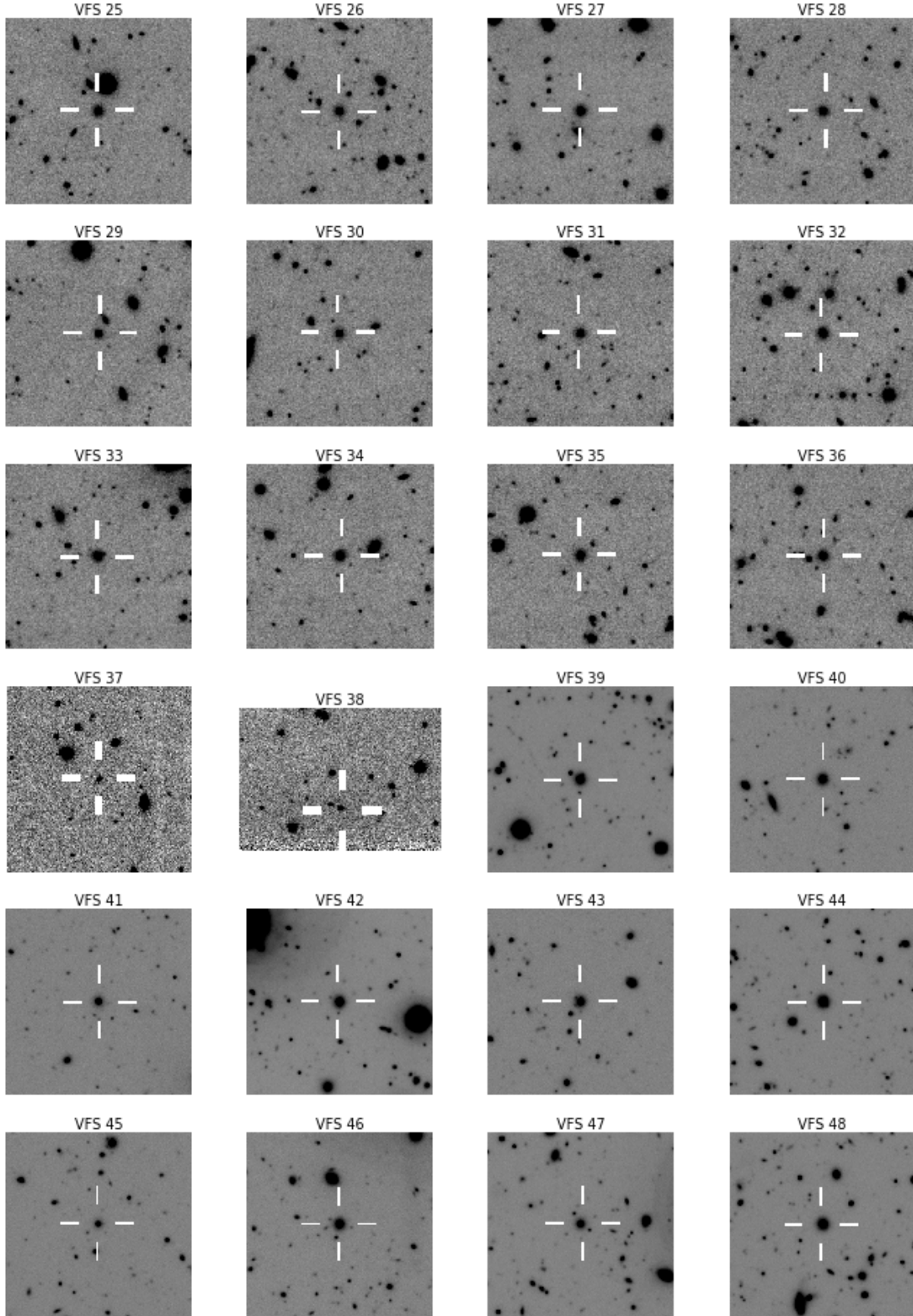


Figure D1. (cont.) J-band finding charts, 1 arcminute field-of-view, for the standards presented above in Table 7. North is up and east to the left.

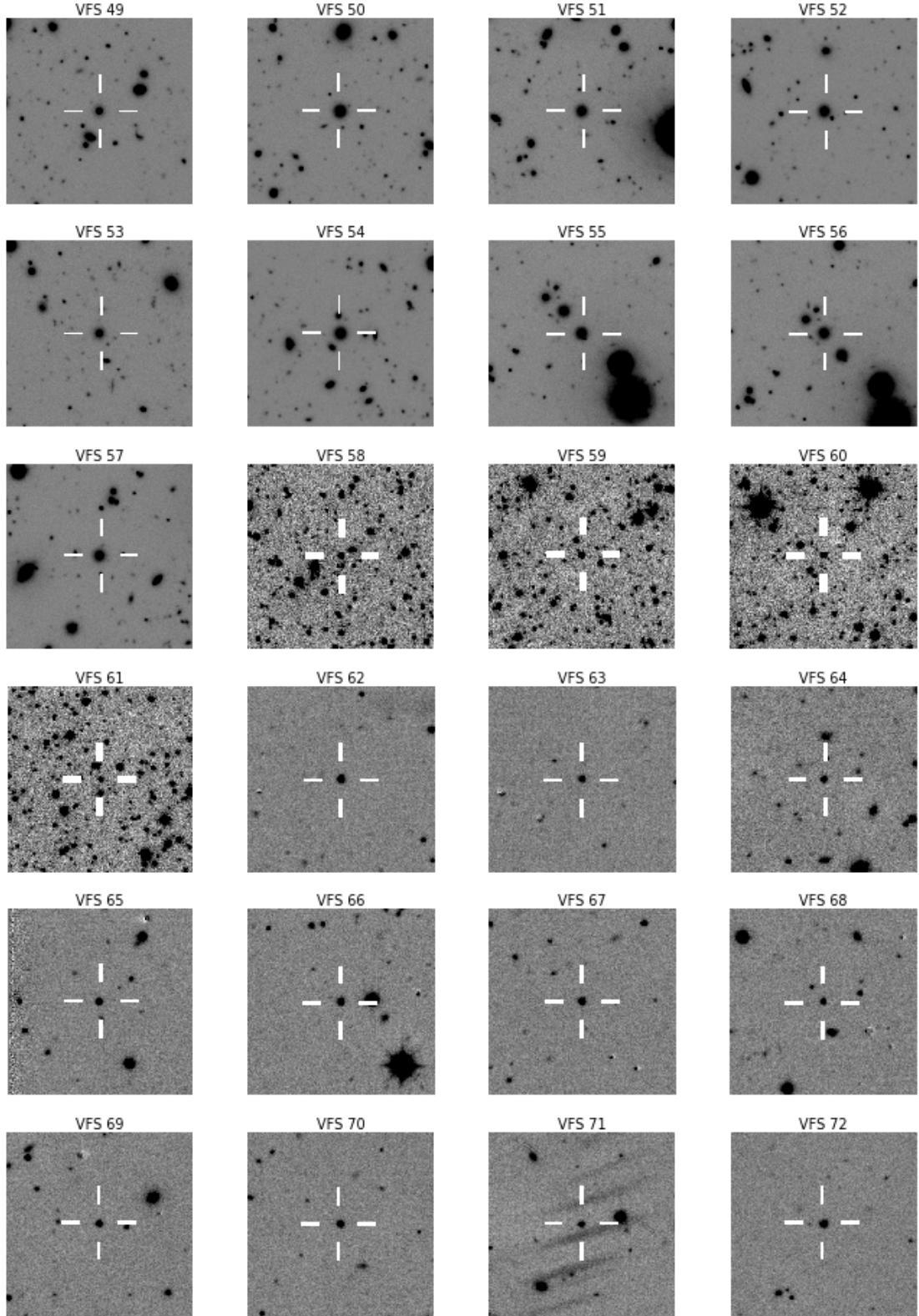


Figure D1. (cont.) J-band finding charts, 1 arcminute field-of-view, for the standards presented above in Table 7. North is up and east to the left.

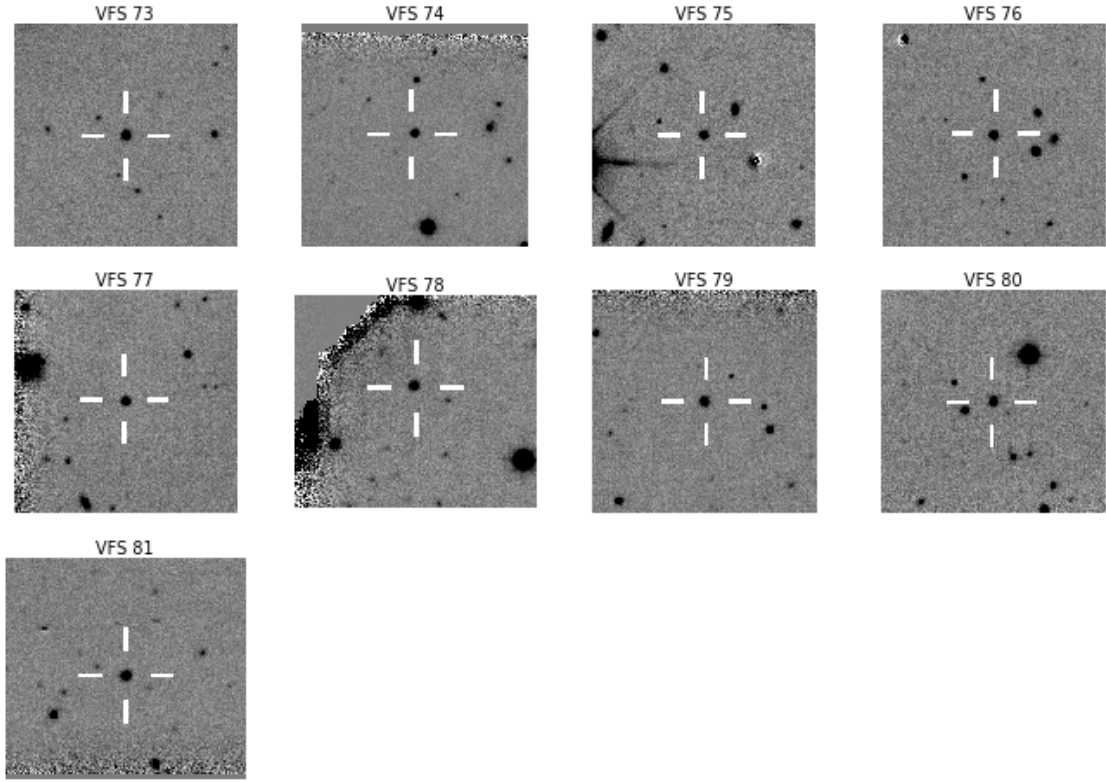


Figure D1. (cont.) J-band finding charts, 1 arcminute field-of-view, for the standards presented above in Table 7. North is up and east to the left.